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# **STEM Imaging**

## **Lecture 17**

**Much of the material for this class is courtesy**

**Nigel Browning of UC Davis & LLNL**

**and**

**Dave Muller of Cornell**

# Outline

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**How does it work?**

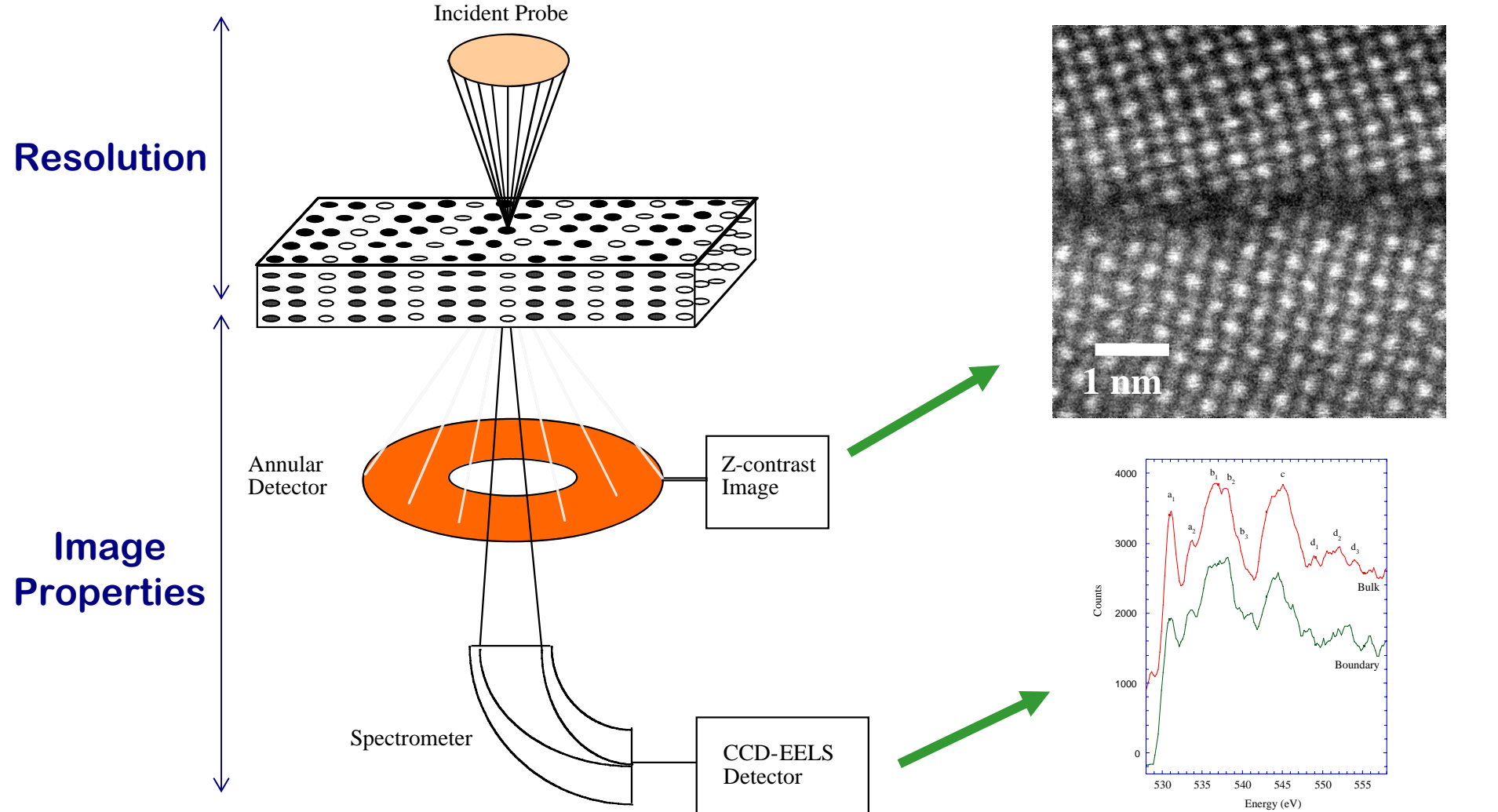
**Inelastic scattering (some review, some new)**

**Instrumental and alignment concerns**

**Image artifacts**

**Examples spread throughout**

# Operational principle

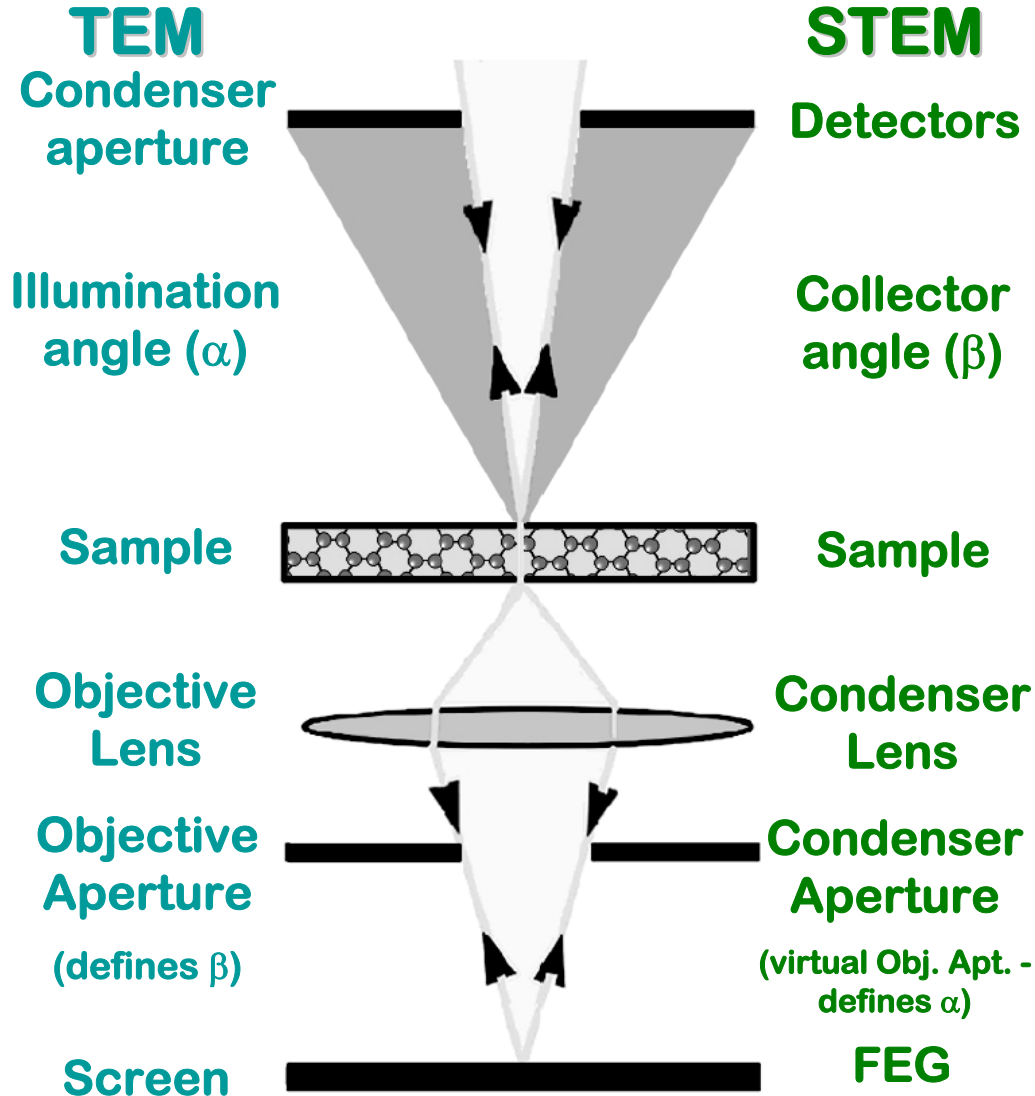




# Reciprocity

In certain instances, TEM & STEM images are strictly equivalent

“Theorem of Reciprocity”



	TEM	STEM
Coherent illumination	Small Condenser Aperture	Small Collector Aperture
Incoherent illumination	Hollow cone illumination	HAADF

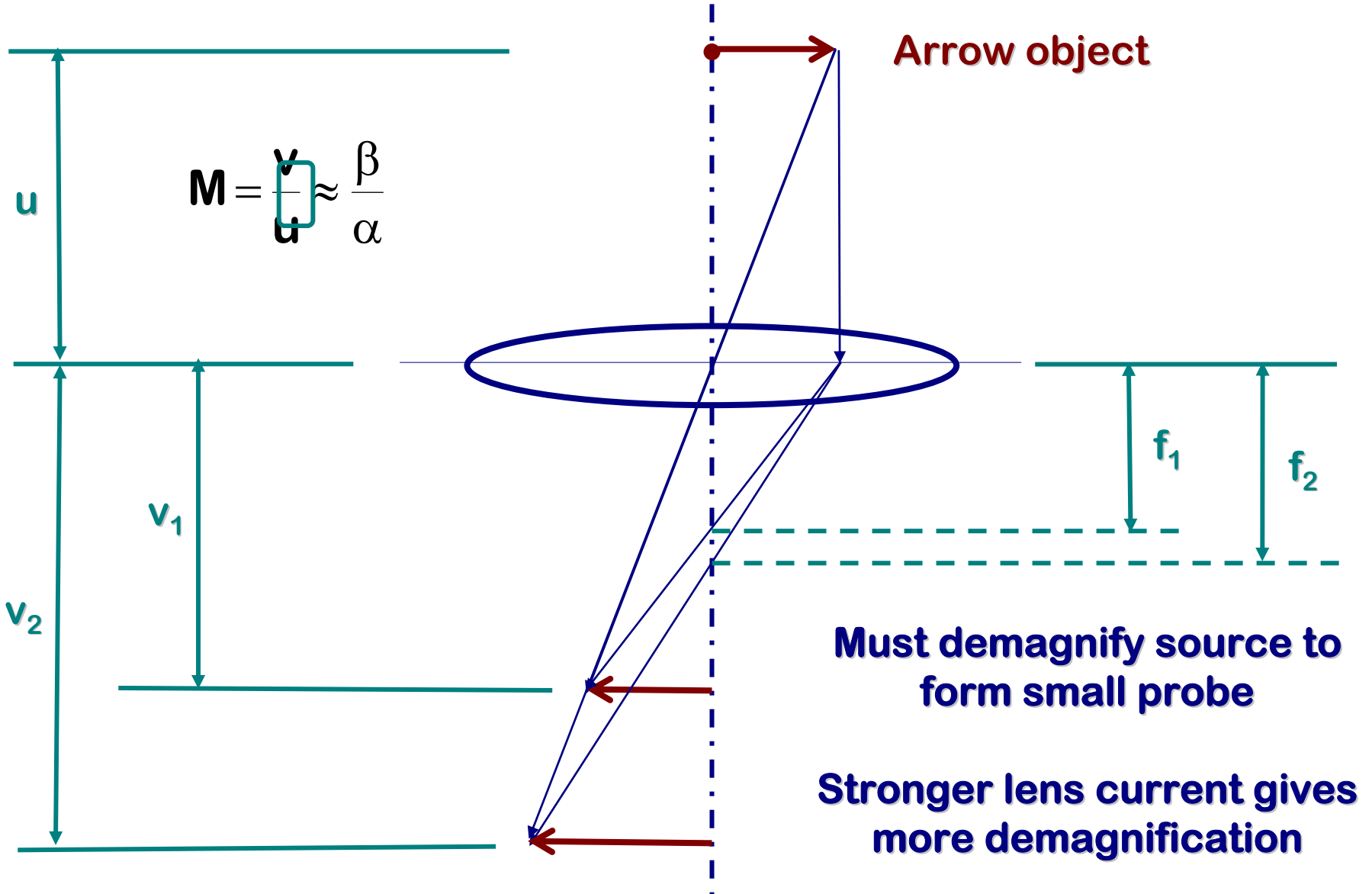


# Electron sources

## A small bright source is necessary for STEM

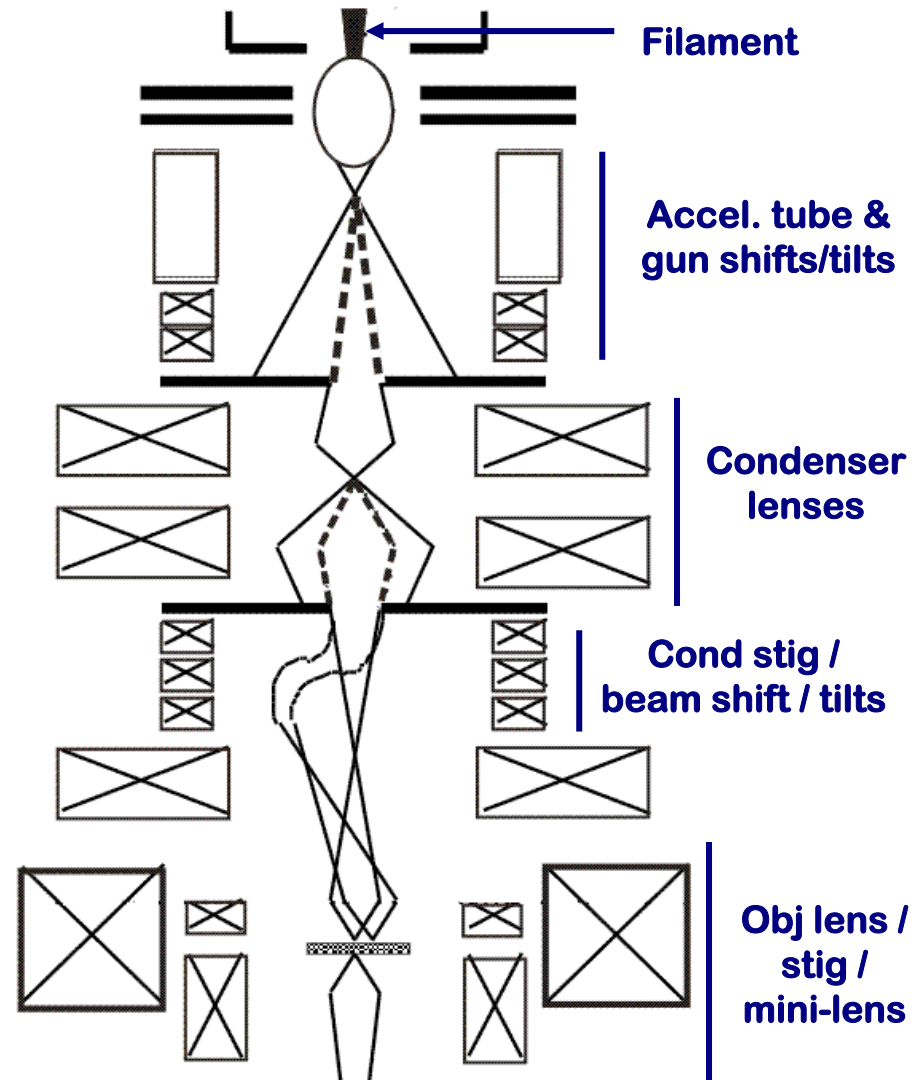
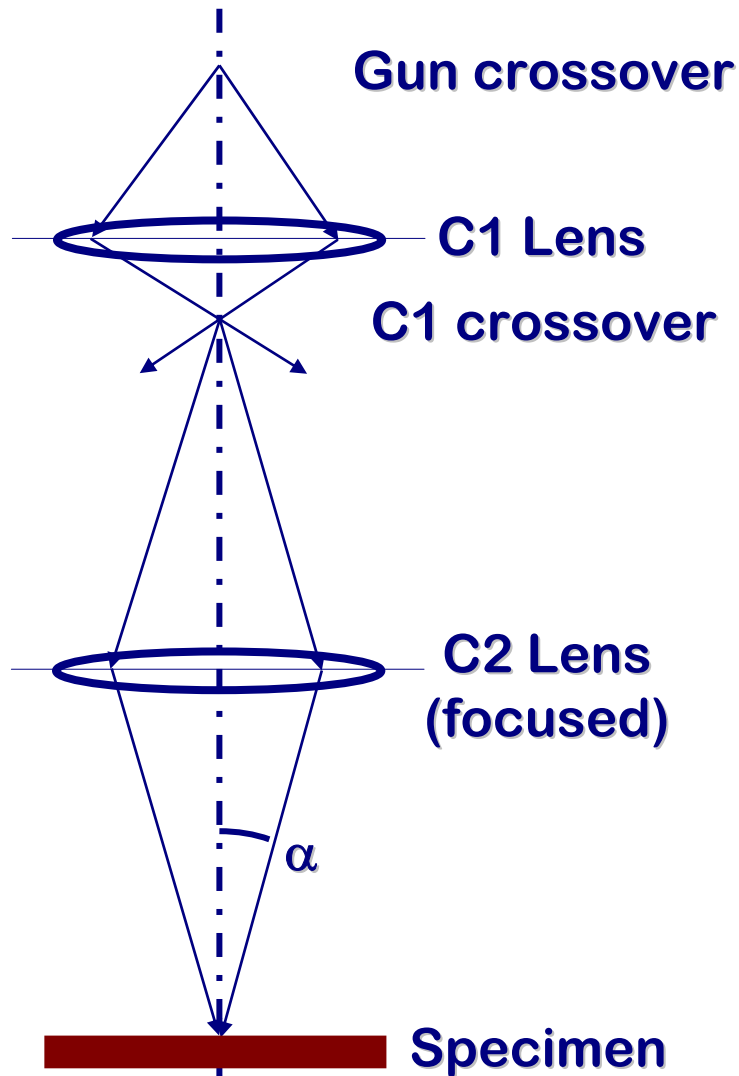
		Thermal emission		Field emission			
		W	LaB <sub>6</sub>	Shottky ZrO/W	Thermal FE W (100)	Cold FE W (310)	
<b>Brightness (A/cm<sup>2</sup>/sr) at 200kV</b>		~5x10 <sup>5</sup>	~5x10 <sup>6</sup>	~5x10 <sup>8</sup>	~5x10 <sup>8</sup>	~5x10 <sup>8</sup>	
<b>Electron Source Size</b>		50μm	10μm	0.1-1μm	10-100nm	10-100nm	
<b>Energy Width (ev)</b>		2.3	1.5	0.6-0.8	0.6-0.8	0.3-0.5	
<b>Operating Conditions</b>	<b>Vacuum (Pa)</b>	10 <sup>-3</sup>	10 <sup>-5</sup>	10 <sup>-7</sup>	10 <sup>-7</sup>	10 <sup>-8</sup>	10 <sup>-9</sup>
	<b>Temperature (K)</b>	2800	1800	1800	1600	300	
<b>Emission</b>	<b>Current (μA)</b>	~100	~20	~100	20-100	5-20	
	<b>Short term stability</b>	1%	1%	1%	7%	5%	2%
	<b>Long term stability</b>	1%/hr	3%/hr	1%/hr	6%/hr	20%	10%
<b>Maintenance</b>		Not necessary	Not necessary	Start-up takes time	Build up necessary after change	Flash every few hours	
<b>Price &amp; Operation</b>		Low & simple	Low & simple	High & easy	High & easy	High & complicated ?	
<b>Lifetime</b>		3 months	1 year	>4 years	?	?	1 year

# Demagnification





# Probe forming optics



# Scattering

**For perfect crystals, four sources that give high-angle scatter**

- **Rutherford scattering**
  - **Elastic**
  - **Proportional to square of atomic weight ( $\propto Z^2$ )**
- **Higher Order Laue Zone reflections**
  - **Elastic**
  - **Important in thin samples**
- **Thermal diffuse scattering (TDS)**
  - **Scatter from phonons (lattice vibrations)**
  - **Not strictly proportional to  $Z^2$** 
    - **Depends on  $b$  (low angles screened by other planes in lattice)**
    - **Measured between  $Z^{1.5}$  and  $Z^{1.7}$**
- **Electron Compton scattering**
  - **Inelastic scatter off of the electrons**

# Types of STEM images

## Bright-field

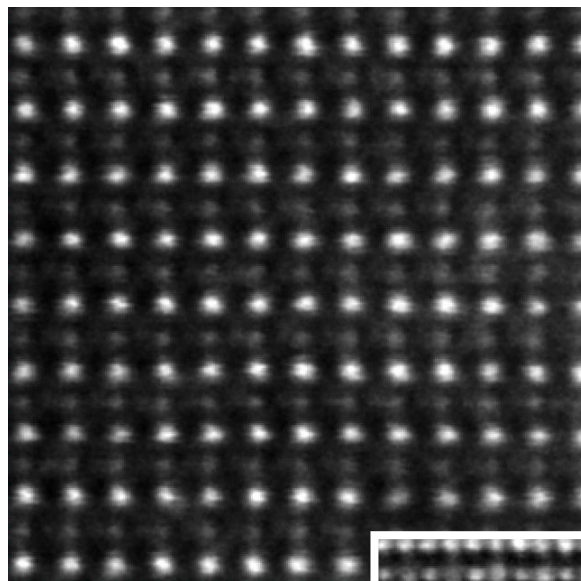
- Collect central beam with a small collection angle
- Contains elastic (Rutherford), phonon, plasmon and Compton

## Low-angle annular dark field

- Collection angle of 25 - 50 milliradians (mrad)
- Mostly phonon scatter

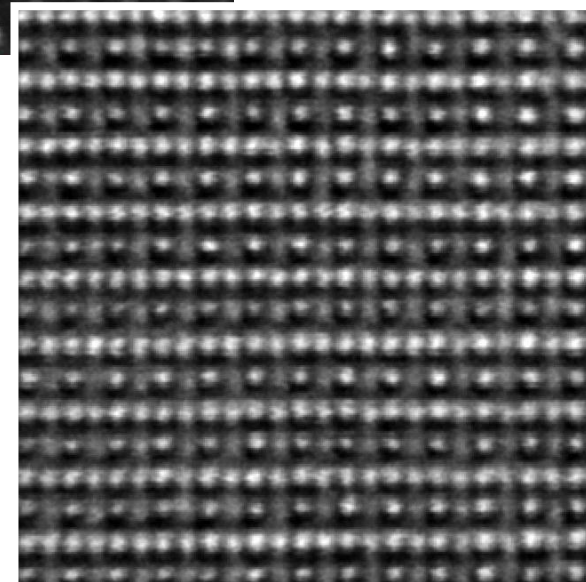
## High-angle annular dark field

- Collection angle of 50 - 250 mrad
- Largely phonon scatter (TDS)

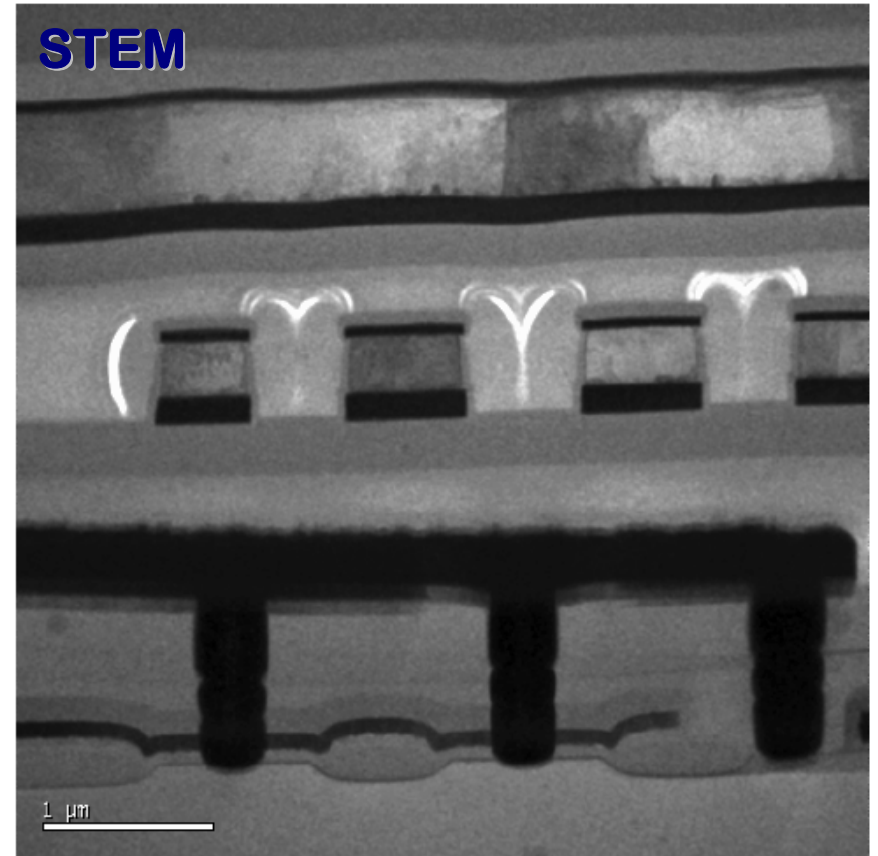
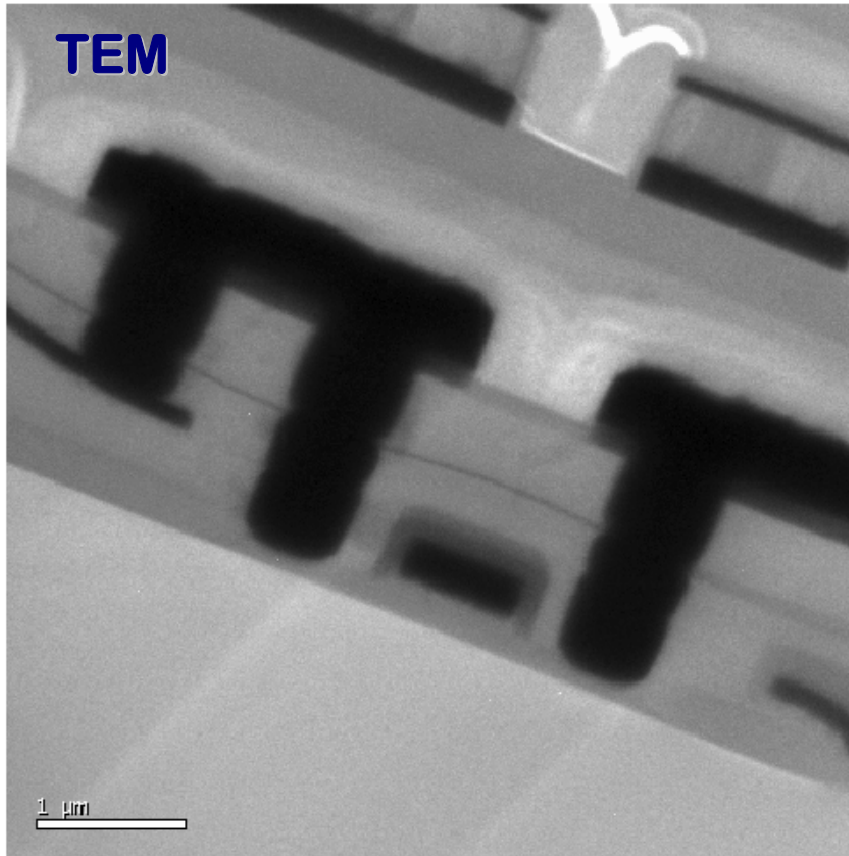


Coherent BF-STEM image of SrTiO<sub>3</sub> <110>

## HAADF-STEM



# BF-STEM

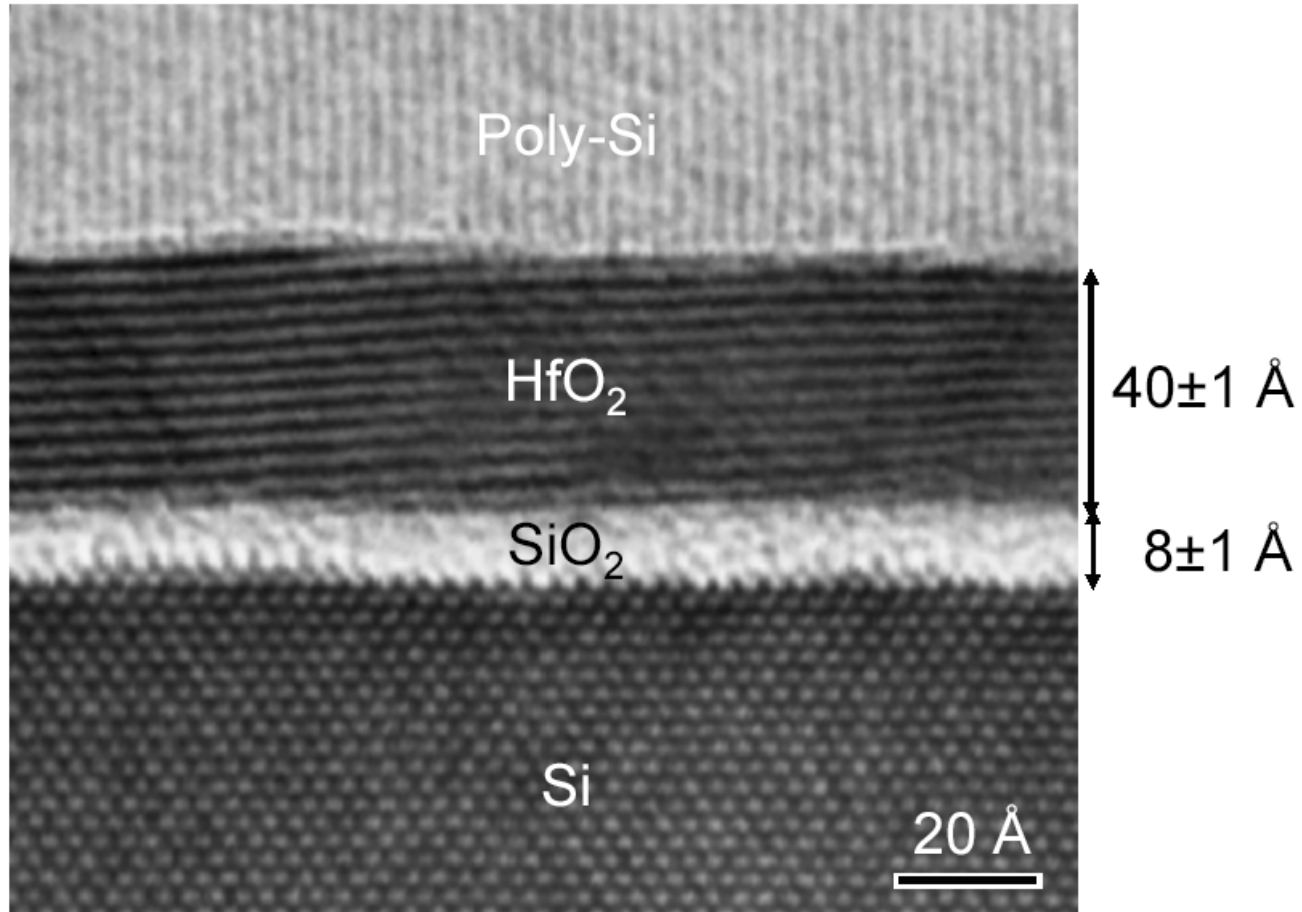


Images courtesy Dave Muller

In STEM, energy losses in sample do not contribute to chromatic aberration (averaged over collector ...)

Finding increasing use in semiconductor quality control

# BF-STEM

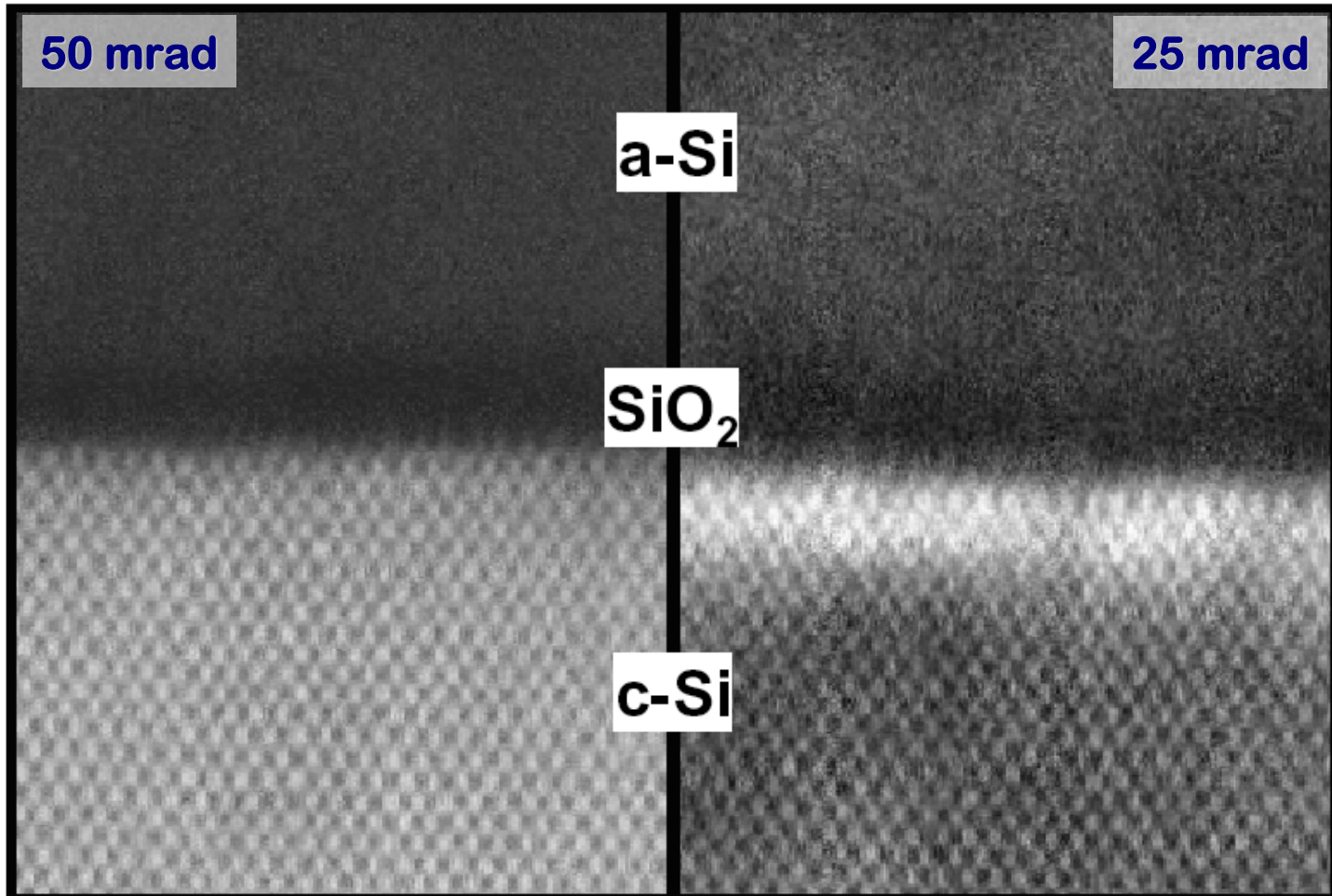


“Reciprocity”

But - HRTEM is certainly quicker, no ‘scan noise’

Images courtesy Dave Muller

# Low-angle annular dark field (LAADF)

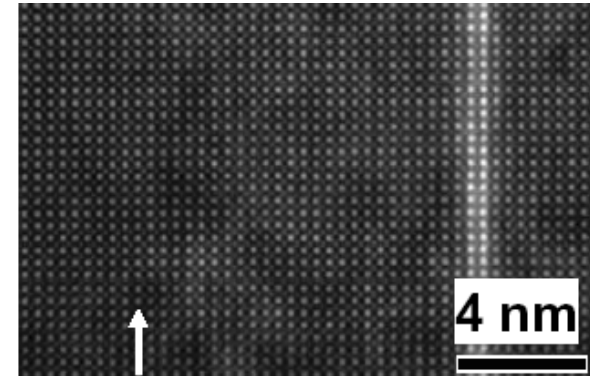


Strain fields cause de-channeling and scattering to small angles

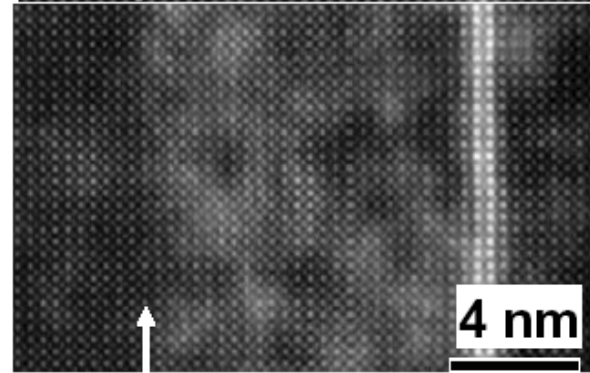
# Low-angle annular dark field (LAADF)

Here contrast is correlated with oxygen vacancies

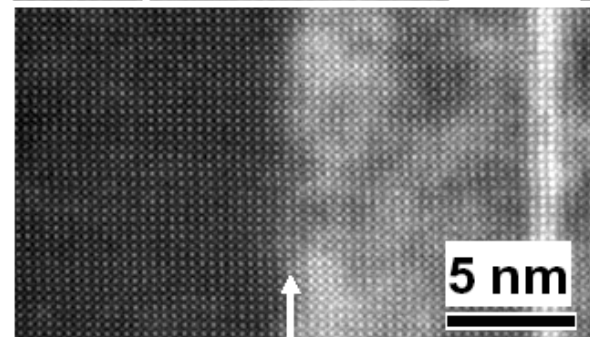
HAADF  
“Z” map



LAADF  
“Strain” map  
Thin x/s



LAADF  
“Strain” map  
Thick x/s

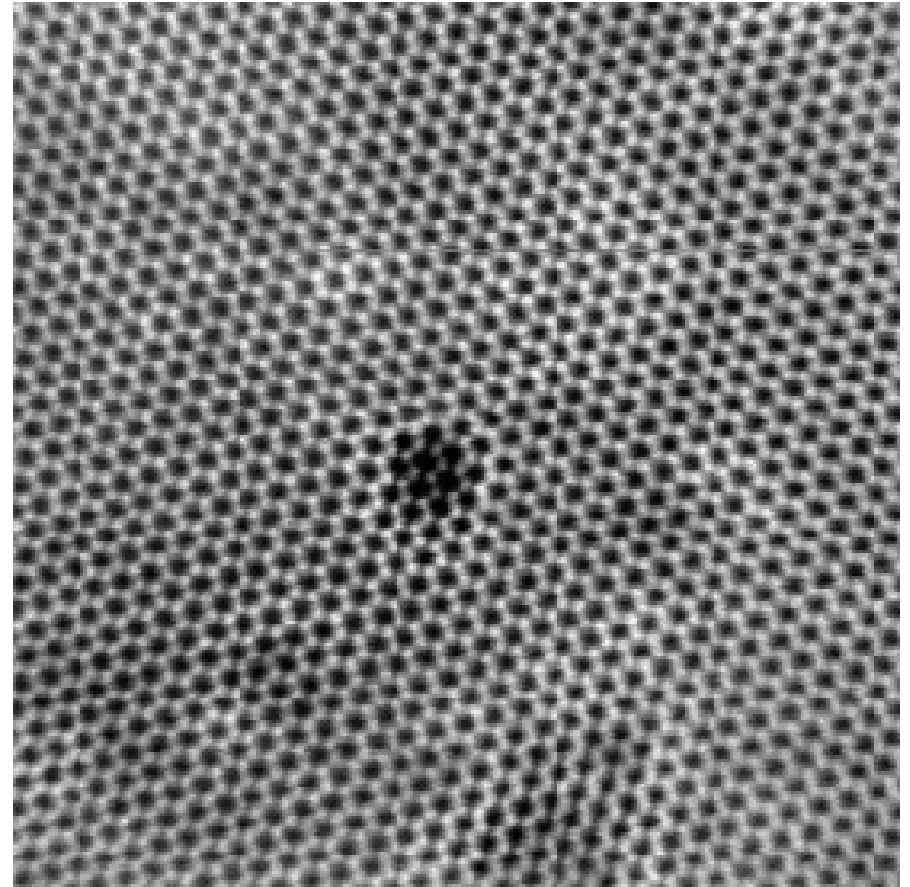


# High angle annular dark field (HAADF)

**No contrast reversals  
with thickness**

**Directly 'interpretable'  
images**

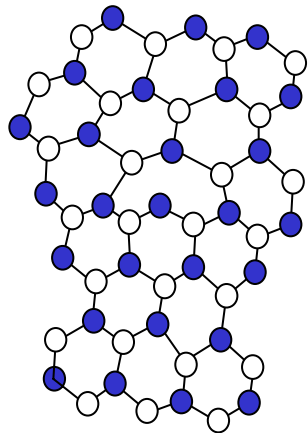
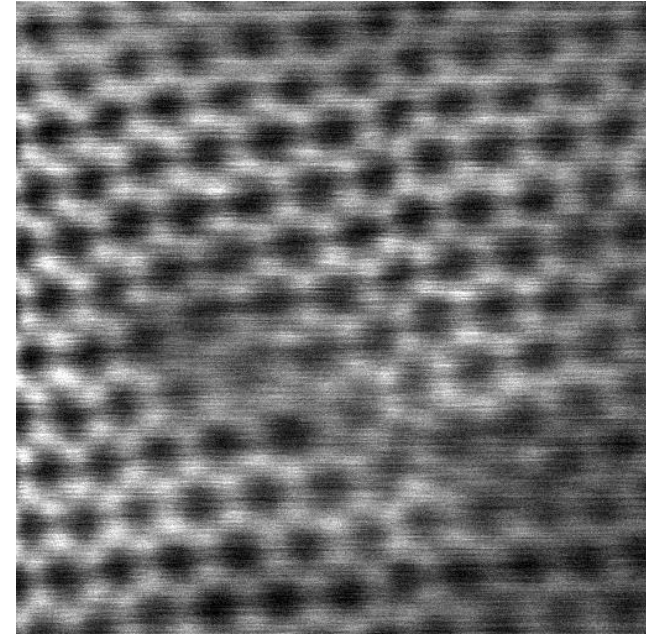
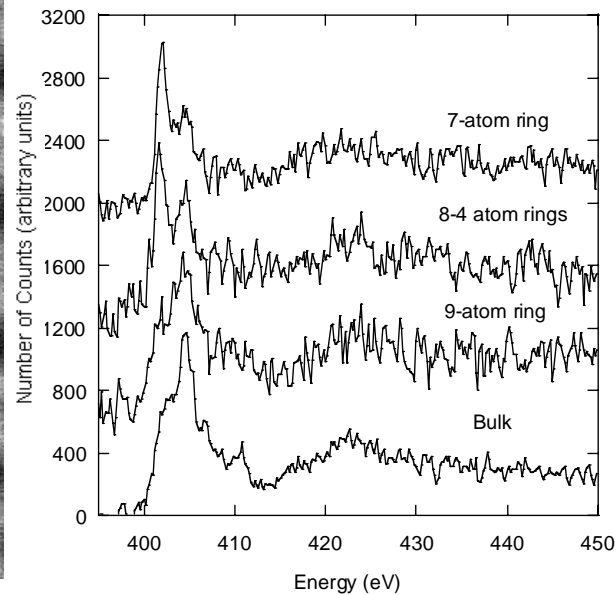
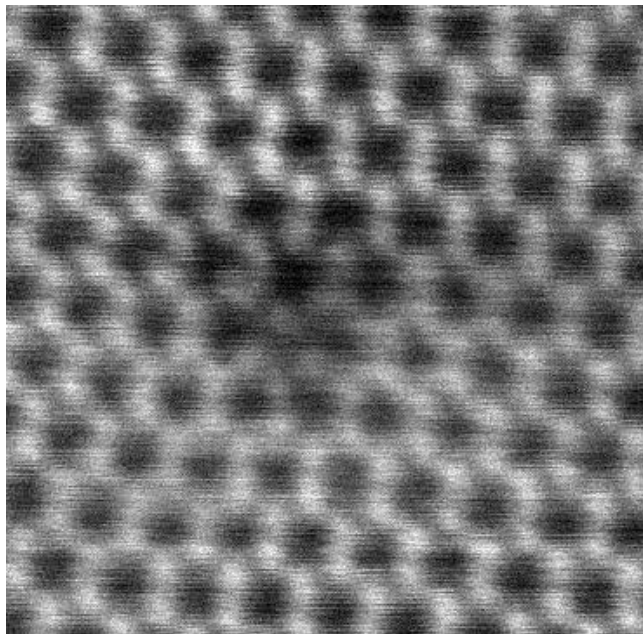
- If you see a white blob, there's an atom column there
- Caveat: the person taking (& processing) the image knew what they were doing ...



**Screw dislocation core in GaN**



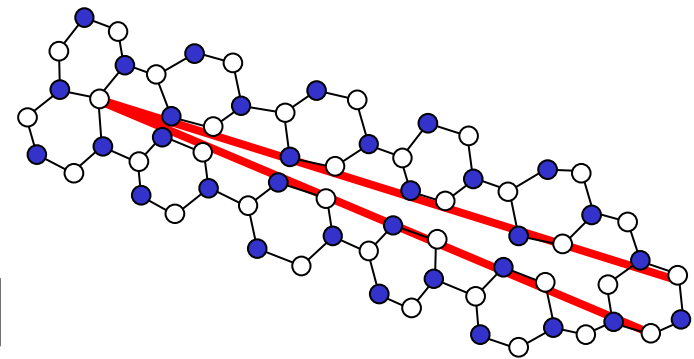
# HAADF of dislocation cores



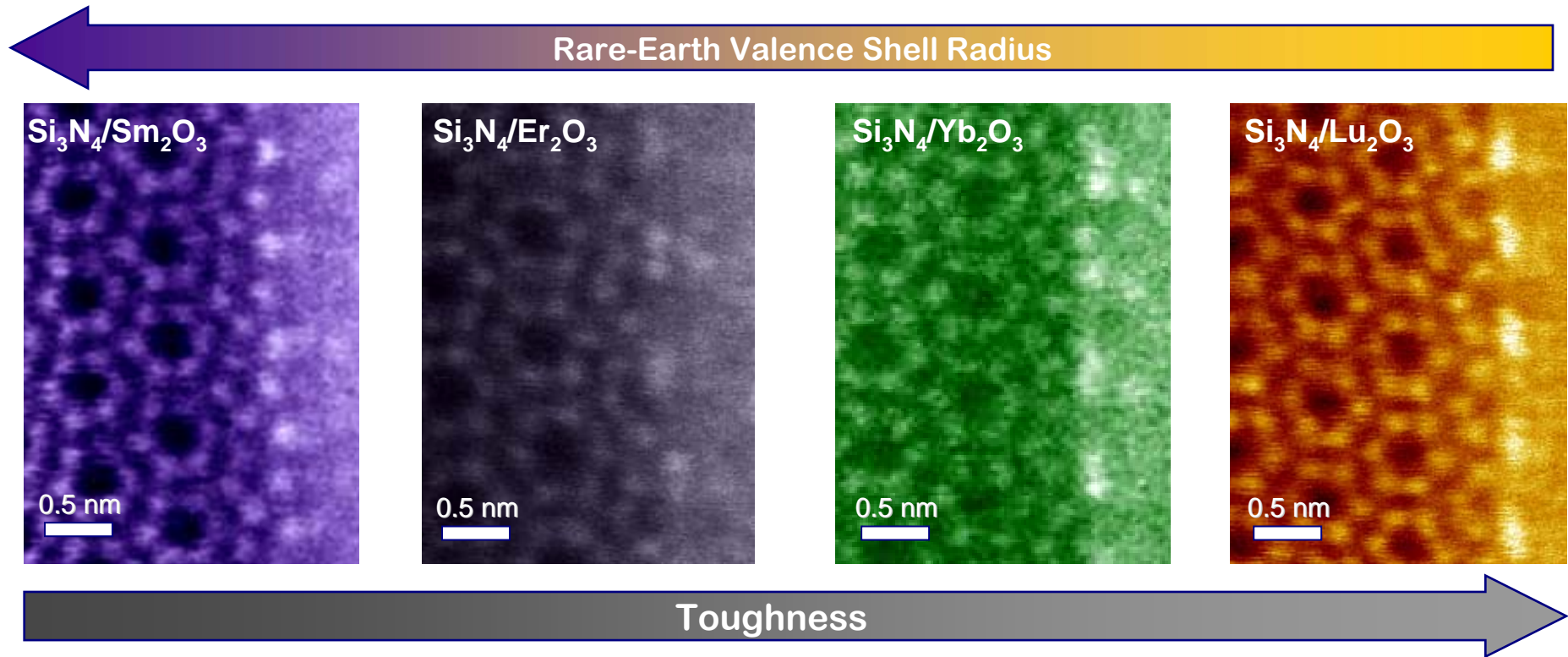
$$\frac{1}{3} [113] \rightarrow \frac{1}{3} [110] + [001]$$

or

$$\frac{1}{3} [11\bar{2}3] \rightarrow \frac{1}{3} [11\bar{2}0] + [0001]$$



# HAADF images



This is in image of a  $\text{Si}_3\text{N}_4$  sample at a grain boundary lined with amorphous glass.

The effect of different dopant types can be seen

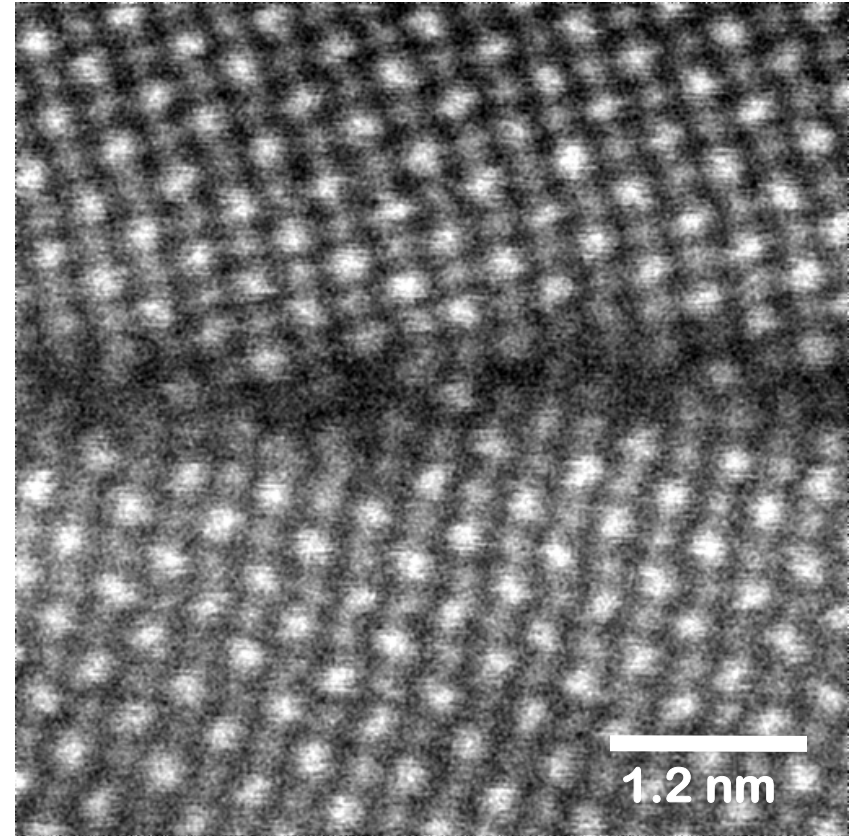
# HAADF of grain boundaries

HAADF often used to study grain boundaries and interfaces

Direct atom positions combined with EELS spectroscopy

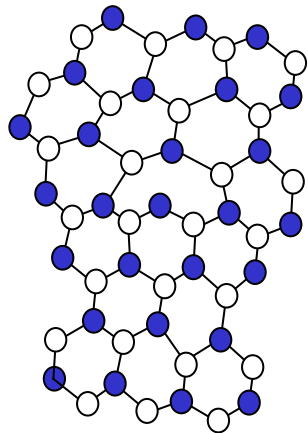
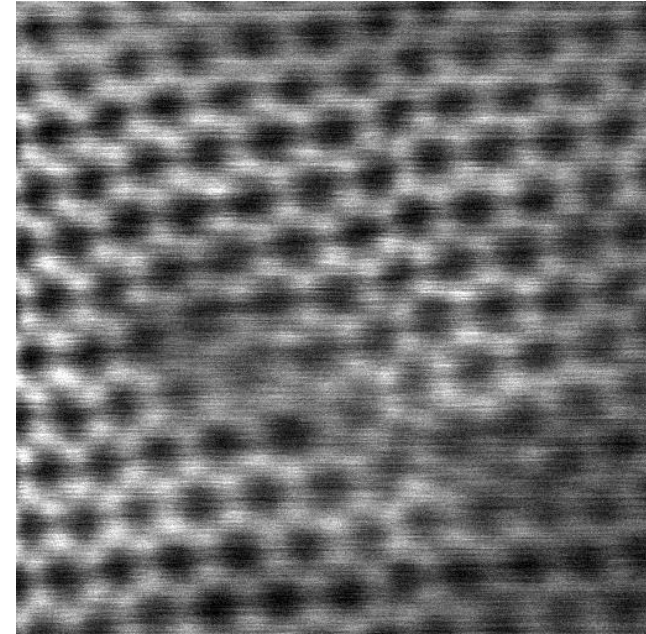
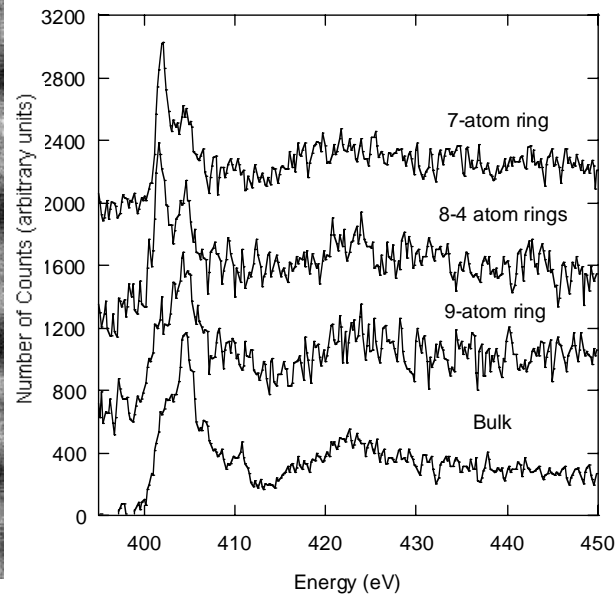
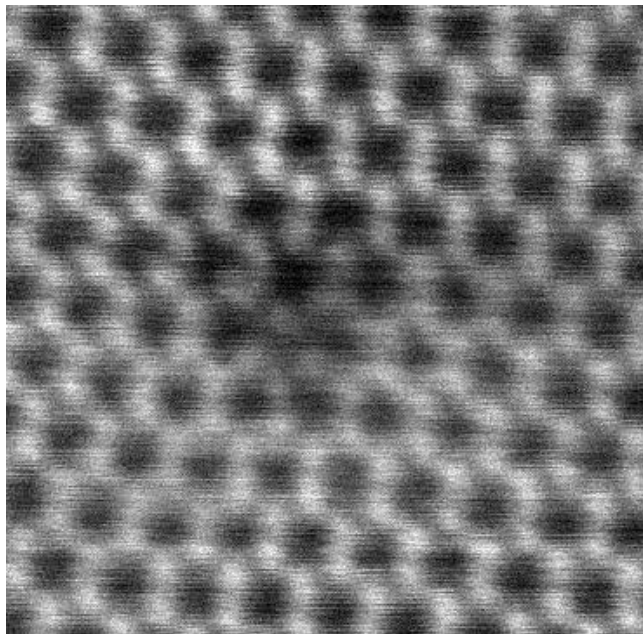
Difference in intensity due to locally different phonon scattering

- Local lattice relaxation allows different phonon modes
- “Huang Scatter”



Grain boundary in SrTiO<sub>3</sub>

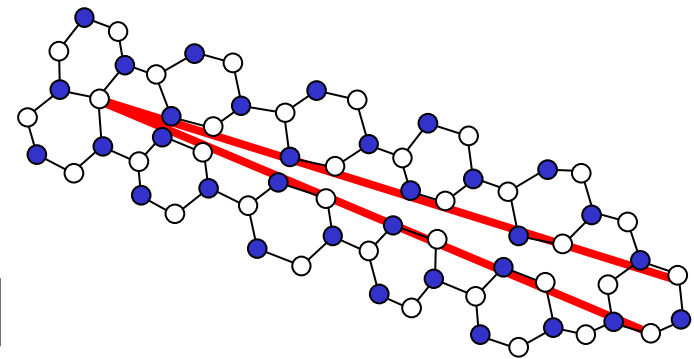
# HAADF of dislocation cores



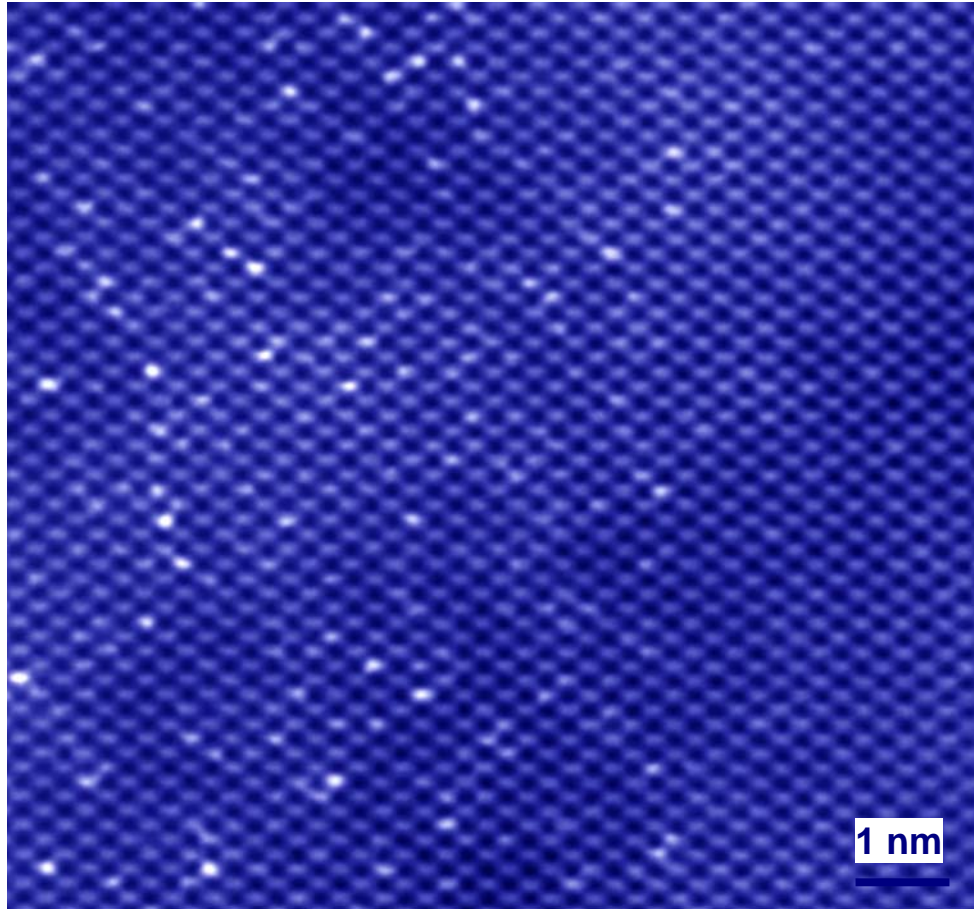
$$\frac{1}{3} [113] \rightarrow \frac{1}{3} [110] + [001]$$

or

$$\frac{1}{3} [11\bar{2}3] \rightarrow \frac{1}{3} [11\bar{2}0] + [0001]$$



# Imaging individual dopant atoms (Sb in Si)



Null test:

No Sb in substrate

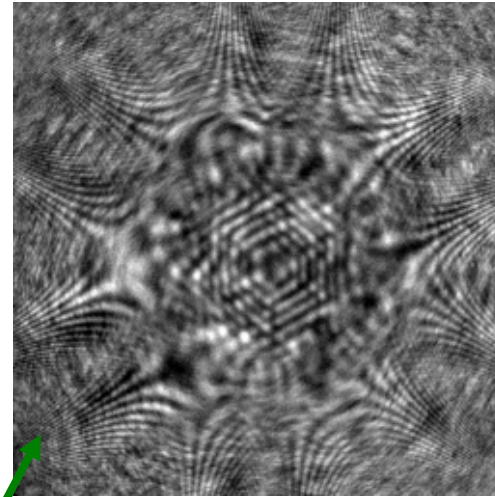
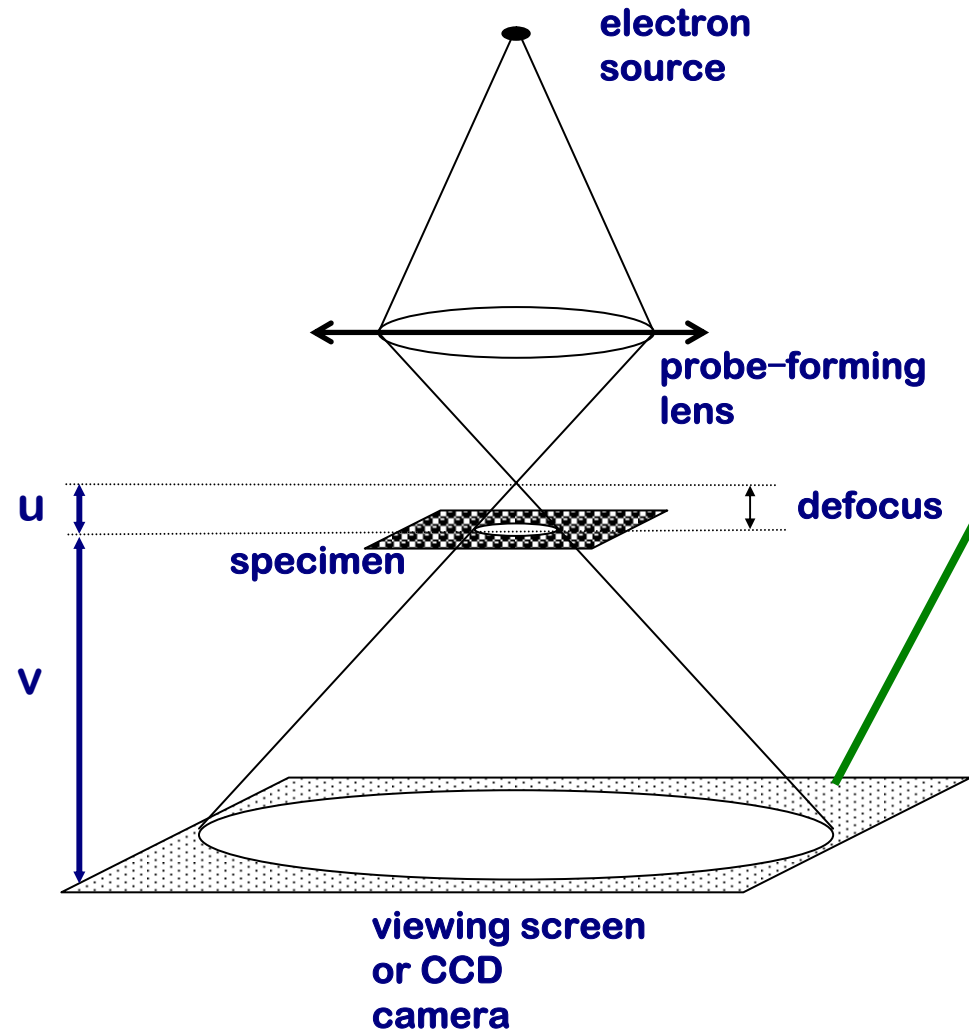
Sb source turned on here

No Sb in substrate

P. Voyles, D. Muller, J. Grazul, P. Citrin, H. Gossmann, *Nature* **416** 826 (2002)

# Alignment

## Electron “Ronchigram”



$$M = \frac{v}{u}$$

**Easiest method of alignment**

**Easiest way to find optic axis**

**Easiest way to correct astigmatism**

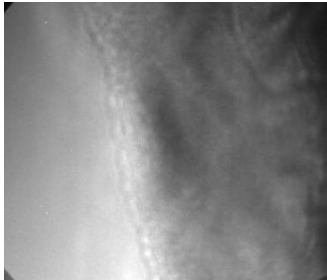
**Easiest way to find focus**

**Works best on amorphous material**

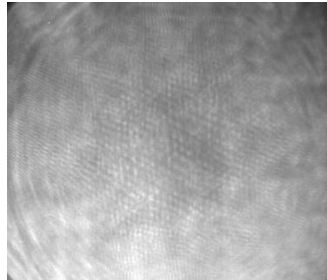
**Start with largest aperture and work your way down**

# Ronchigrams from Si <110>

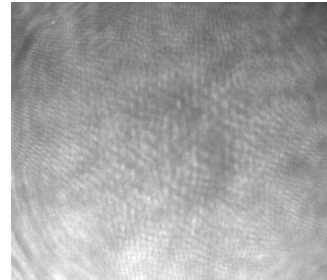
overfocused



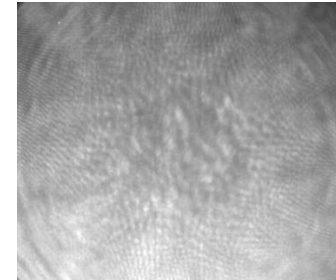
$\Delta f=150$  nm



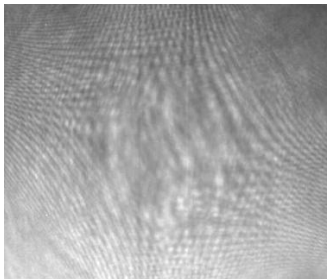
$\Delta f=100$  nm



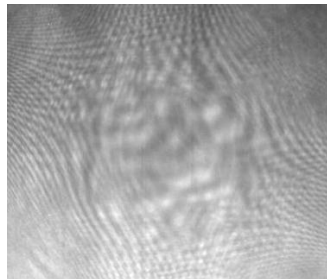
$\Delta f=50$  nm



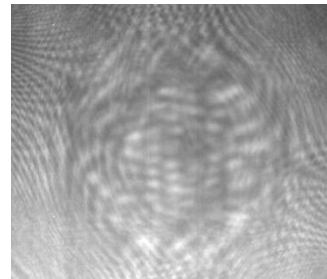
$\Delta f=0$  nm



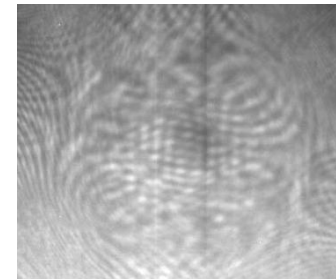
Scherzer focus



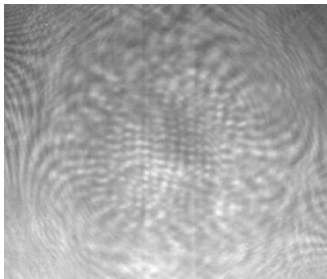
$\Delta f=-100$  nm



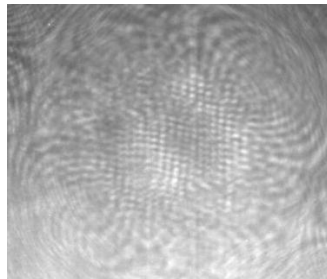
$\Delta f=-150$  nm



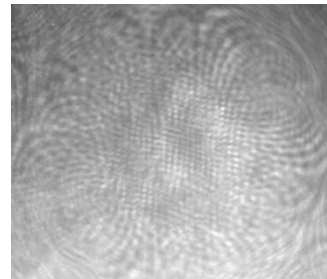
$\Delta f=-200$  nm



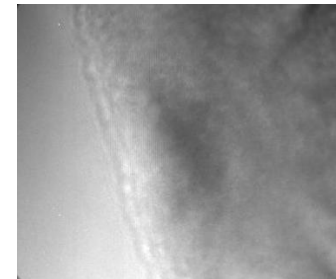
$\Delta f=-300$  nm



$\Delta f=-350$  nm

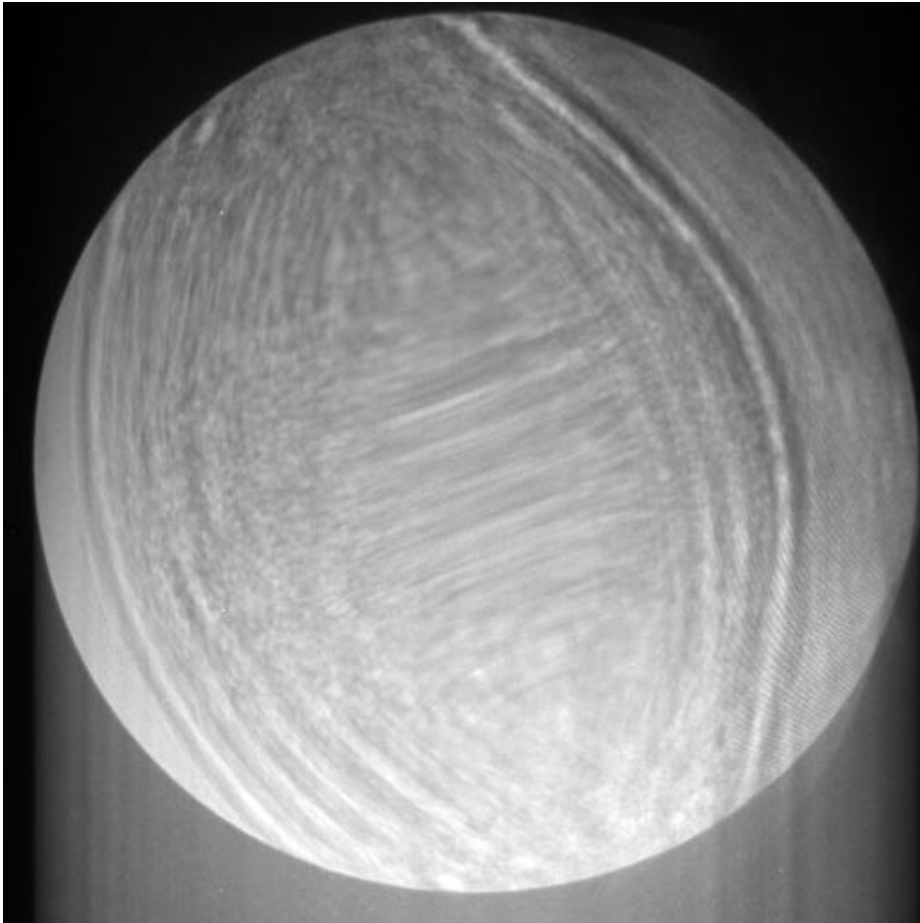


underfocused

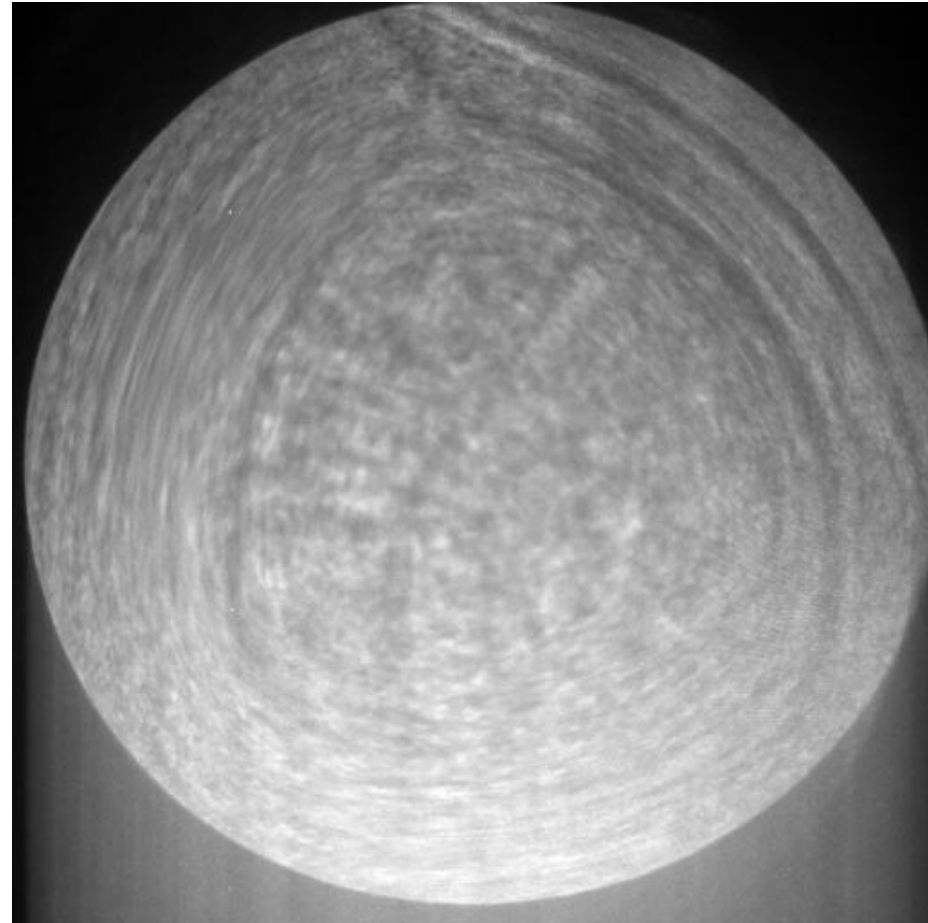


# Correcting for astigmatism

Two fold astigmatism

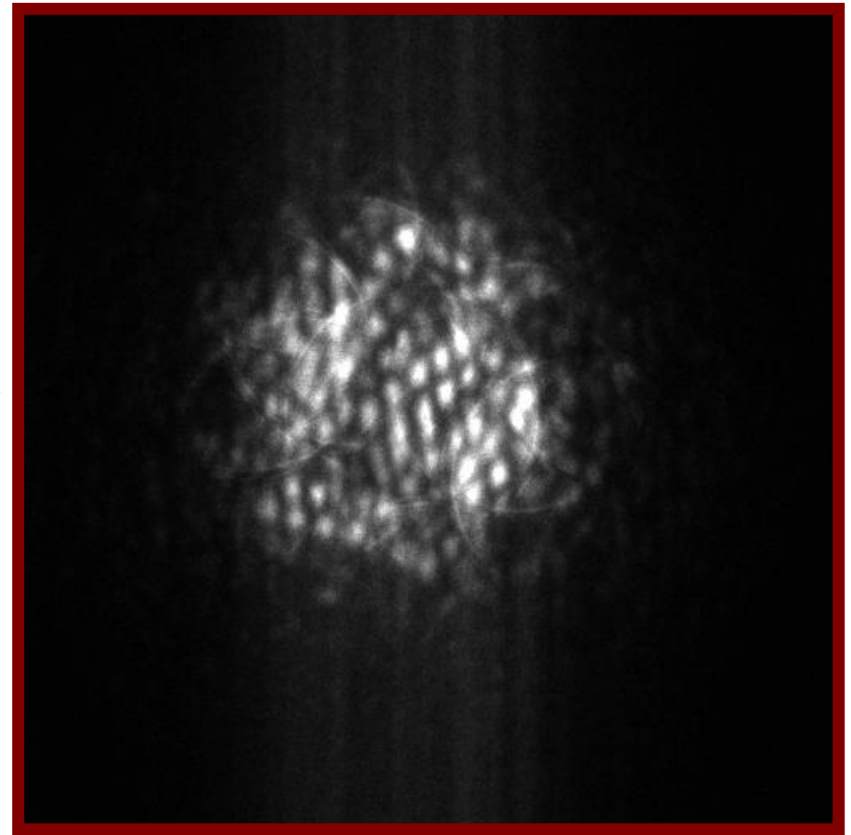
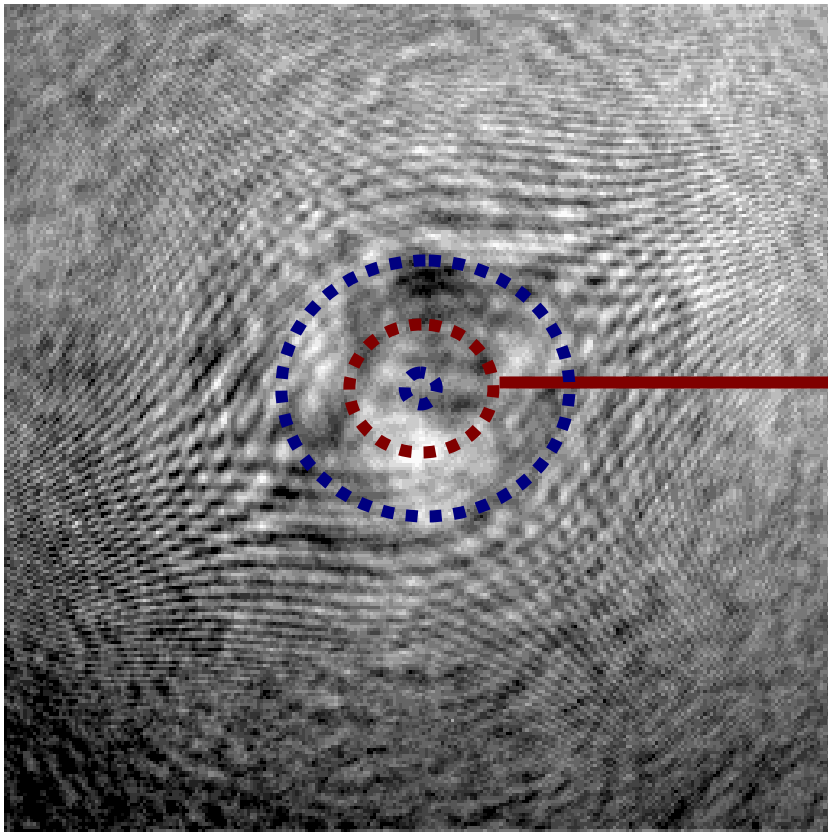


Three fold astigmatism



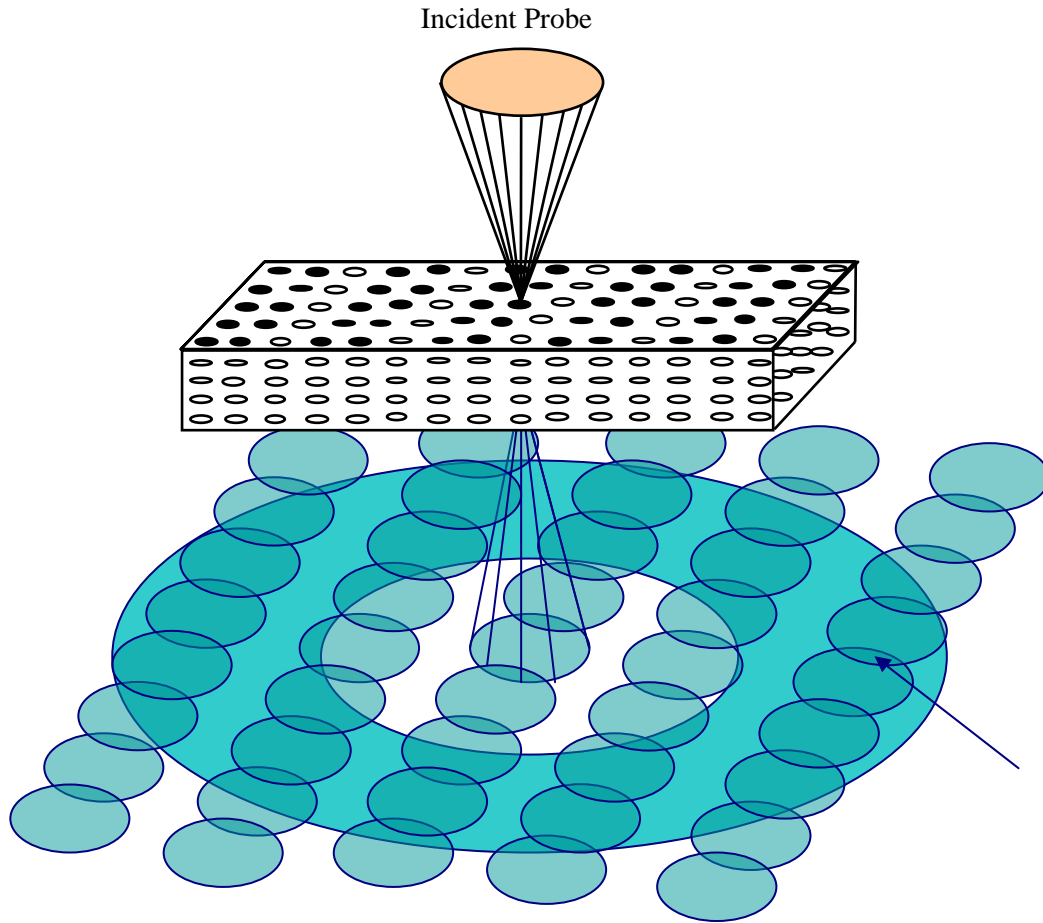


# Forming the Smallest Probe



Put aperture over area of constant phase in Ronchigram to give CBED pattern

# Forming the dark field image



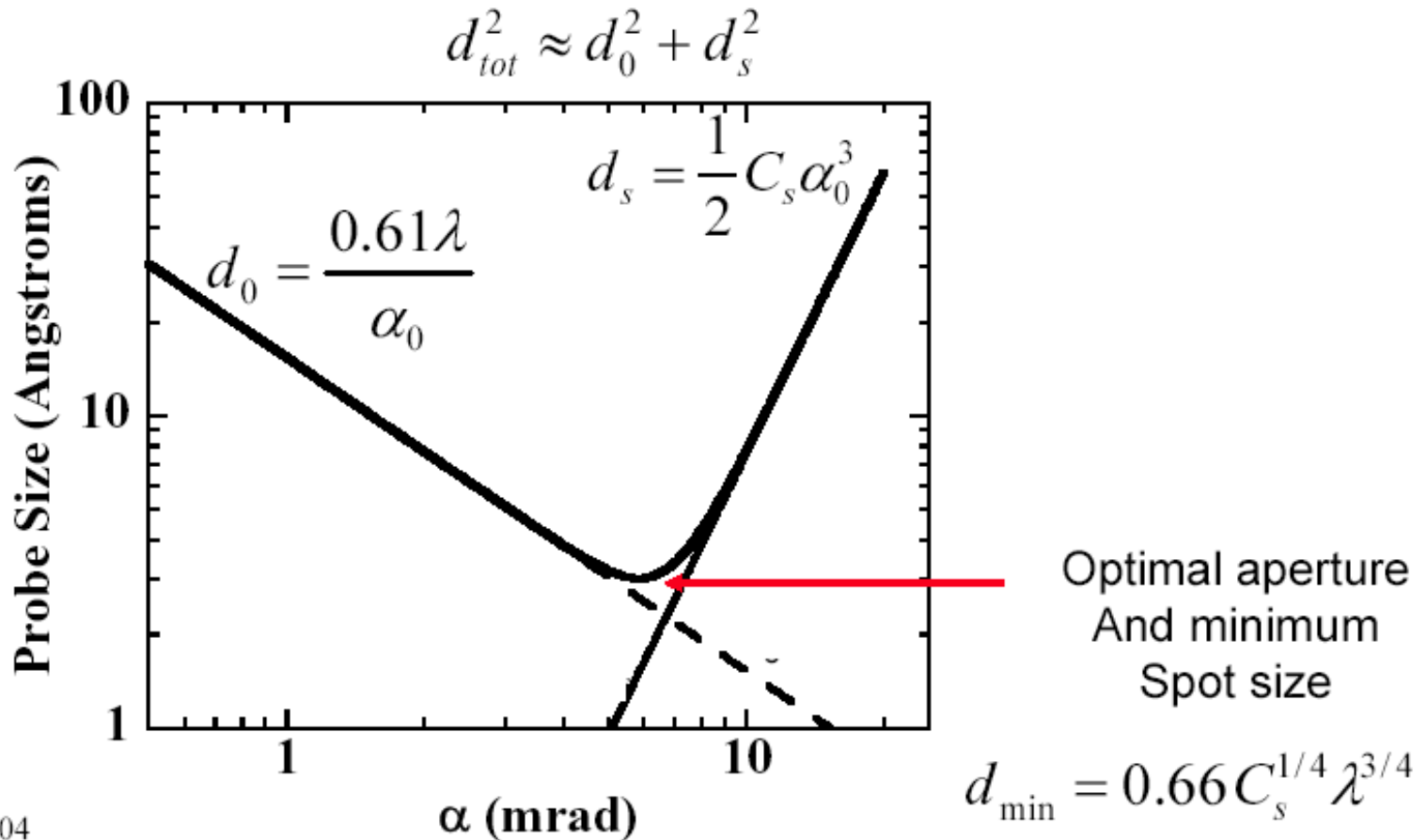
Overlap of disks bright when probe on a column, dark between columns

High-angle detector integrates many spots in the CBED pattern

# Optimum resolution

*balance  $C_s$  and aperture angle*

Remember aperture introduced diffraction effect on resolution



# Optimum resolution

Using wave optics, find:

Minimum spot size:  $d_{\min} = 0.43 C_s^{1/4} \lambda^{1/4}$

Optimum aperture size:  $\alpha_{\text{opt}} = \left( \frac{4\lambda}{C_s} \right)^{1/4}$

At 200kV,  $\lambda = 0.0257\text{\AA}$

$C_s = 1.0 \text{ mm}, d_{\min} = 1.55\text{\AA} \text{ and } \alpha_{\text{opt}} = 10 \text{ mrad}$

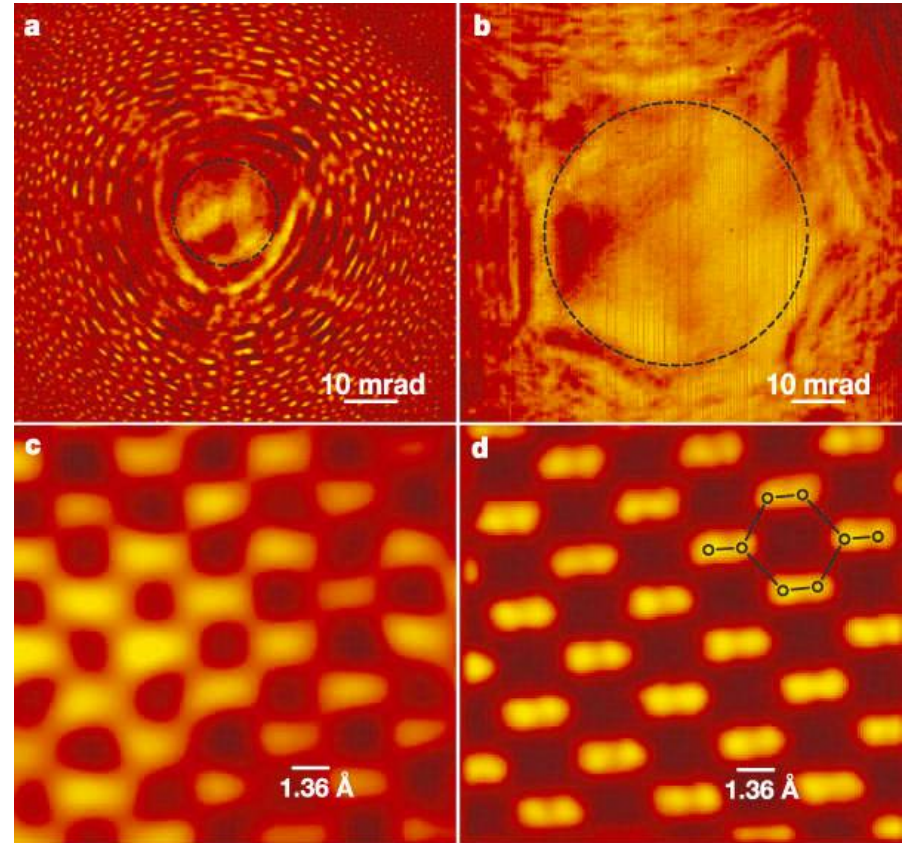
$C_s = 1.2 \text{ mm}, d_{\min} = 1.59\text{\AA} \text{ and } \alpha_{\text{opt}} = 9.6 \text{ mrad}$

$C_s = 0.5 \text{ mm}, d_{\min} = 1.28\text{\AA} \text{ and } \alpha_{\text{opt}} = 12 \text{ mrad}$

$C_s = 0.6 \text{ mm}, d_{\min} = 1.34\text{\AA} \text{ and } \alpha_{\text{opt}} = 11 \text{ mrad}$

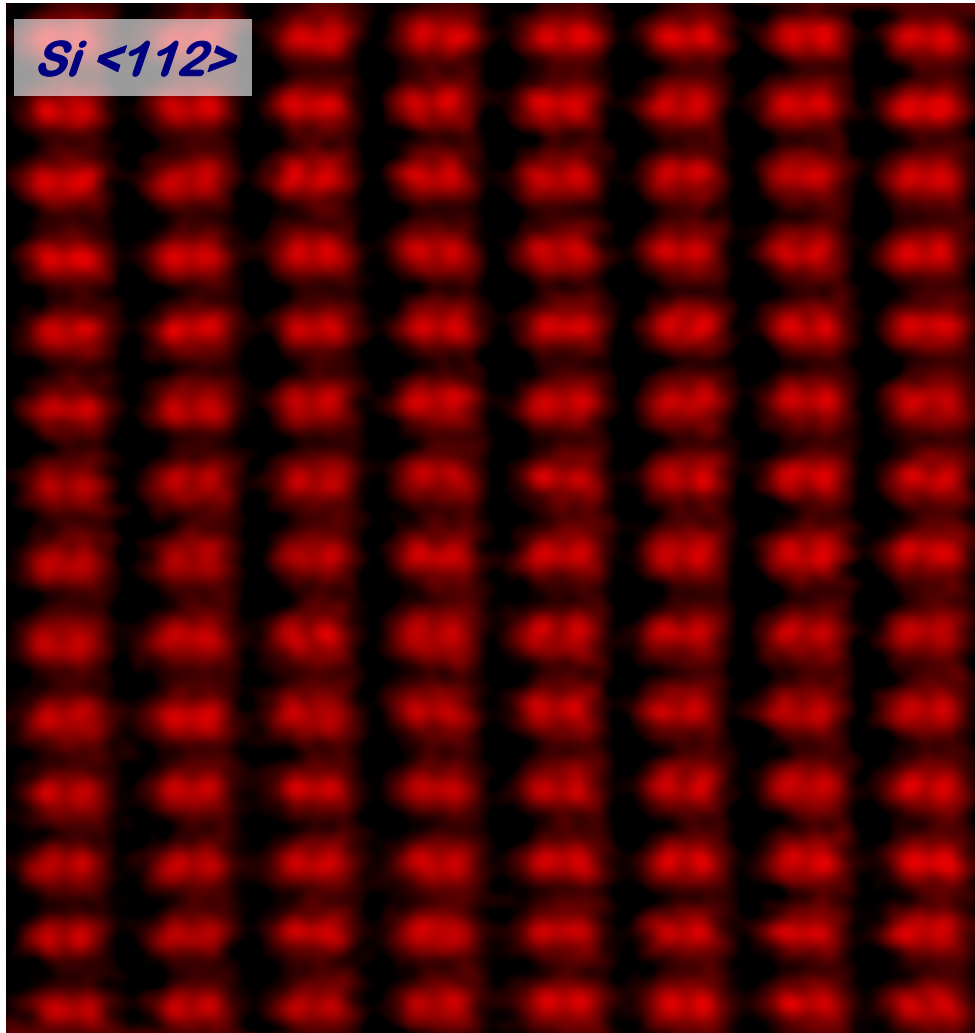
# Effect of $C_s$ Correction

**Cs correction opens up  
the aperture angle  
Allows a much smaller  
probe**



Batson, Dellby and Krivanek, *Nature* 418, 617 (2002)

# Effect of $C_s$ correction



Direct resolution at 0.78 Å

Information transfer to 0.607 Å

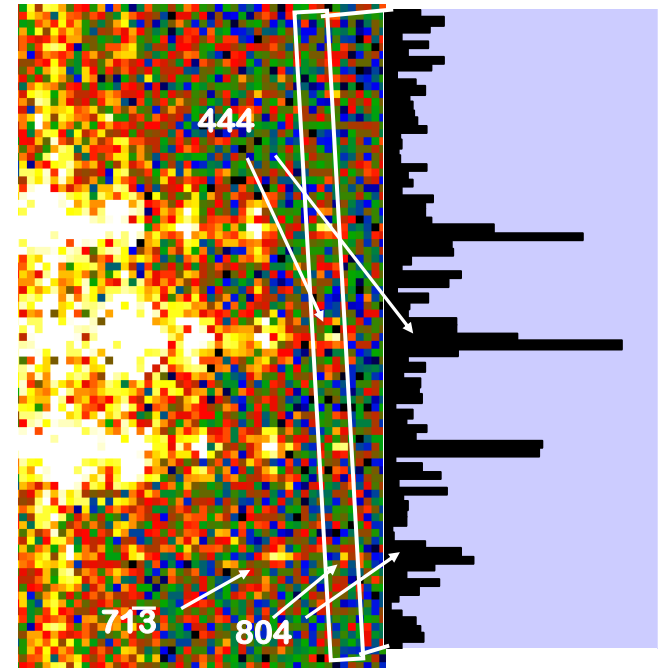


Image by Matt Chisholm,  
Processing by Albina Borisevich and  
Andy Lupini  
Aberration correction by Pete Nellist et  
al., Nion Co.  
Nellist et al., Science 305, 2004.

# Stability issues for STEM

*(and for TEM!)*

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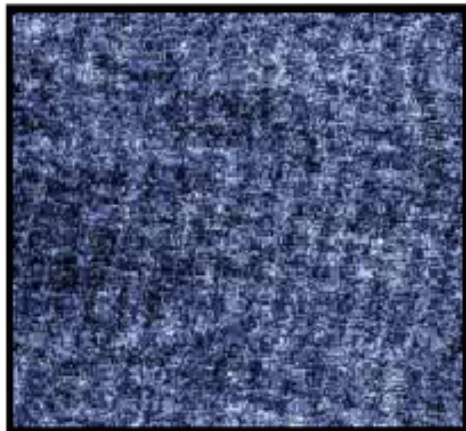
Stray Fields/Ground loops

Air Flow/Temperature Control

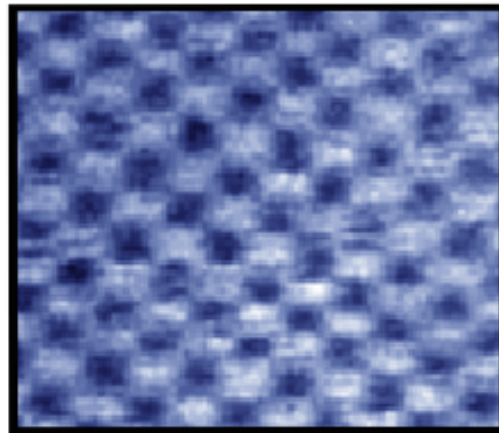
Pressure Variations

Mechanical Vibrations

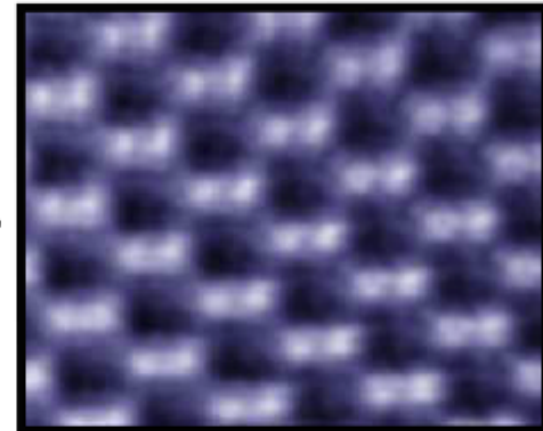
Specimen Preparation



3 weeks

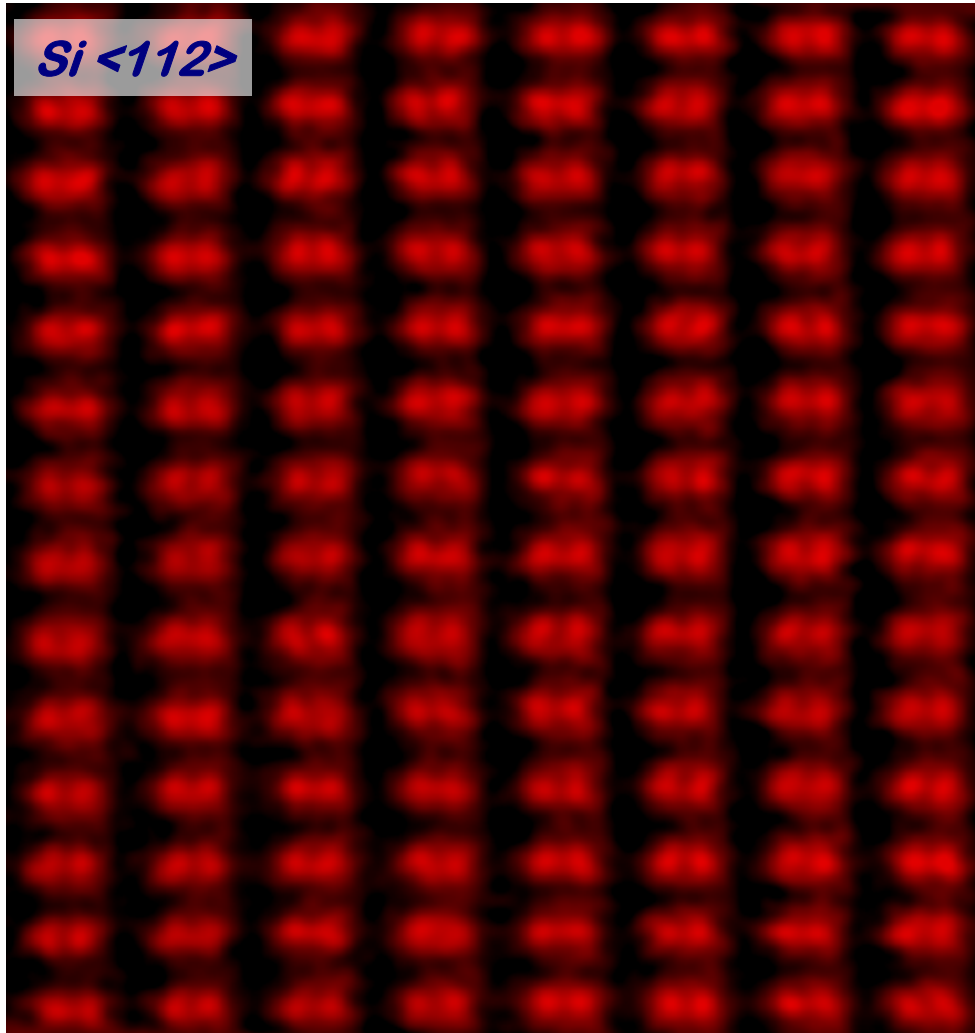


6 months



1 year

# Effect of $C_s$ correction



Direct resolution at 0.78 Å

Information transfer to 0.607 Å

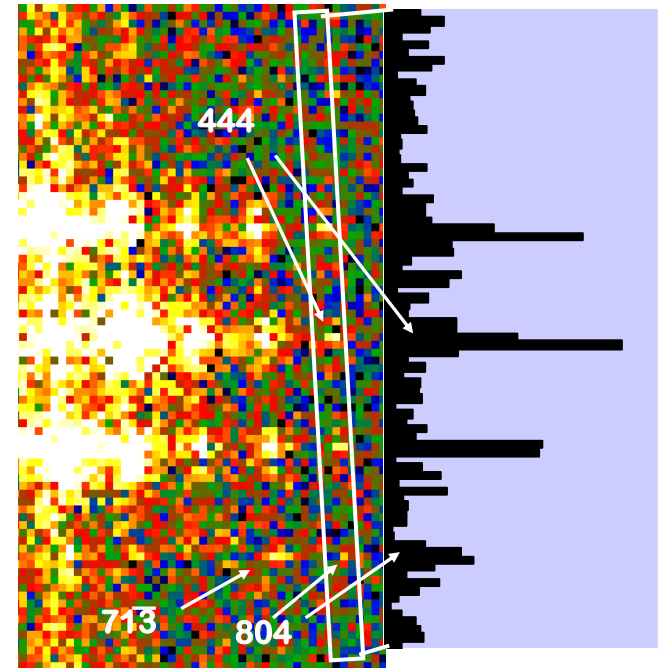


Image by Matt Chisholm,  
Processing by Albina Borisevich and  
Andy Lupini  
Aberration correction by Pete Nellist et  
al., Nion Co.

Nellist et al., Science 305, 2004.



# Image 'artifacts'

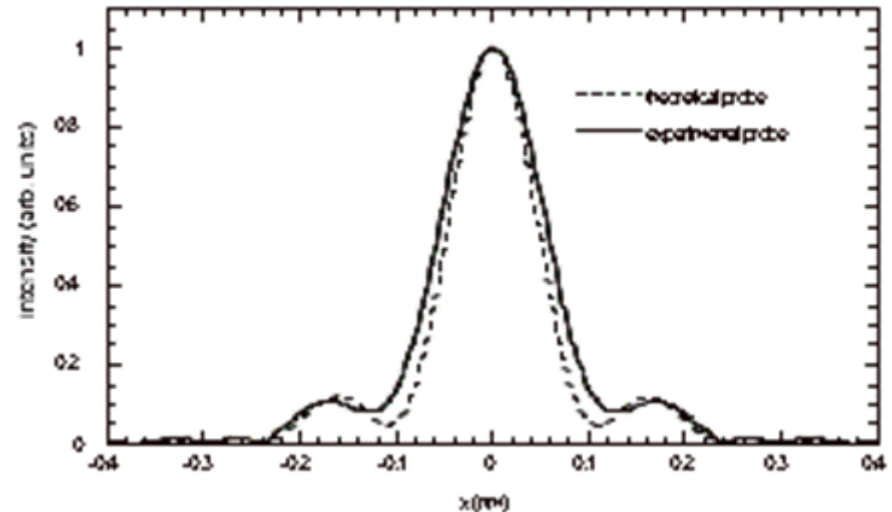
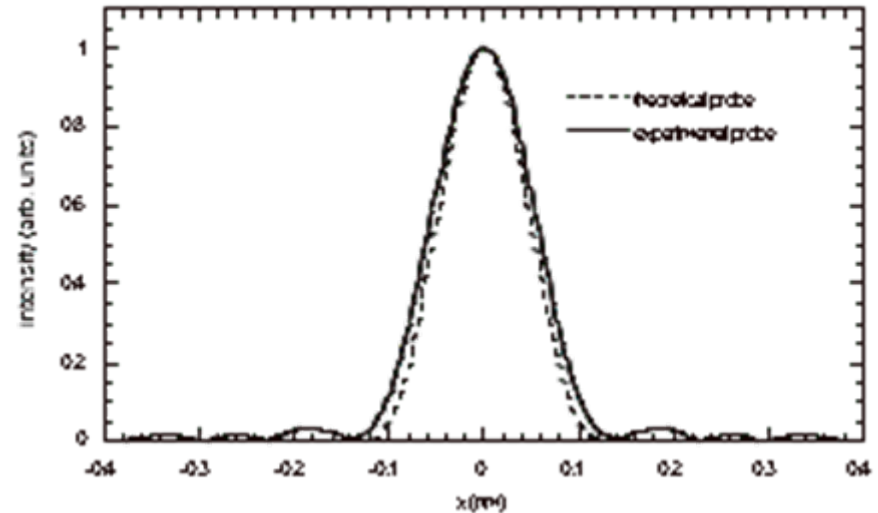
## "probe tails"

Can tweak probe shape to get a narrower probe

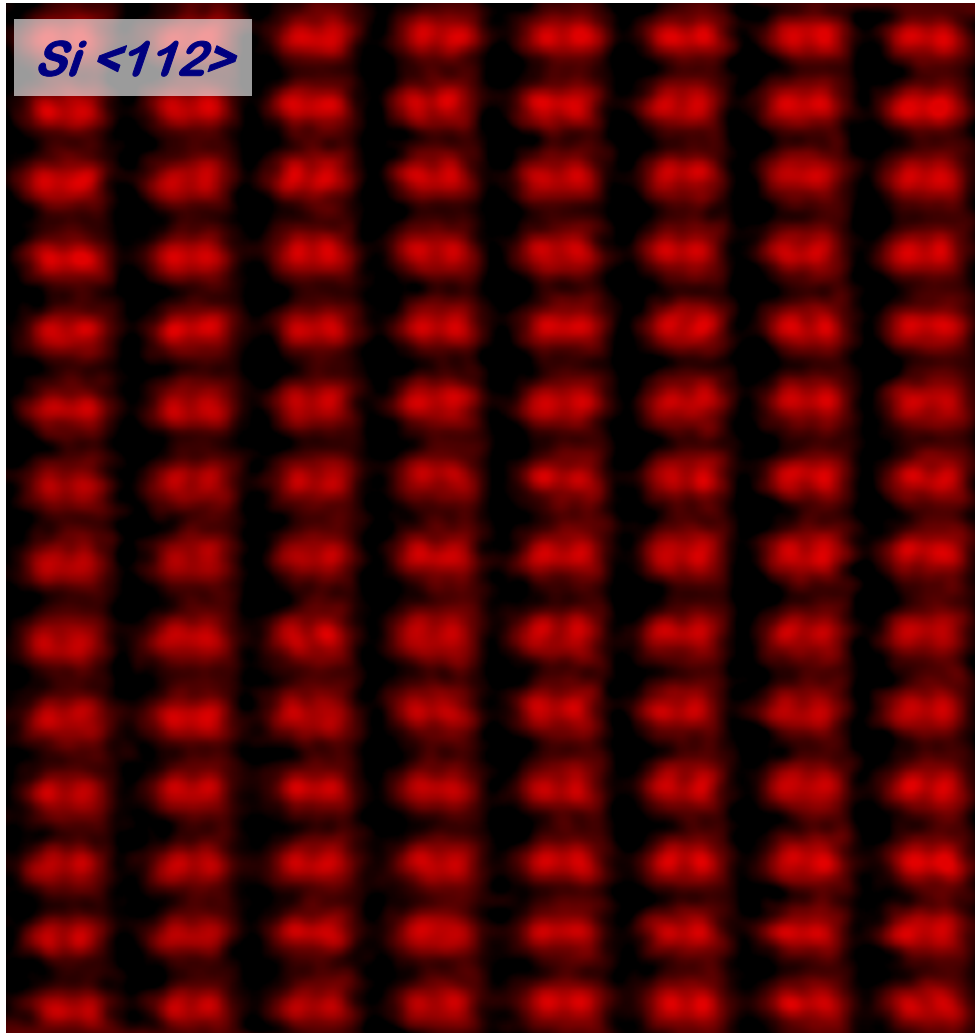
But, gives 'tails' in the distribution

These give 'extra spots' in the diffraction pattern

- The presence of extra spots can be confuse resolution determination
- Now people used images along known projections



# Effect of $C_s$ correction



Direct resolution at 0.78 Å

Information transfer to 0.607 Å

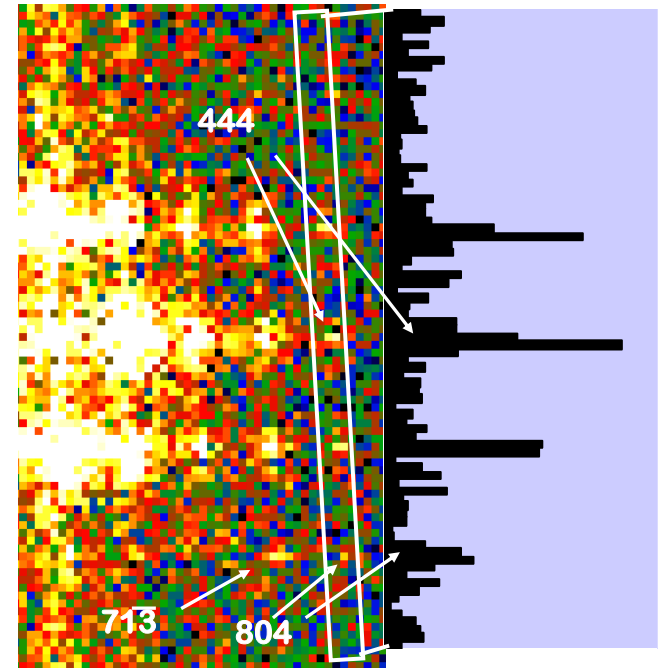
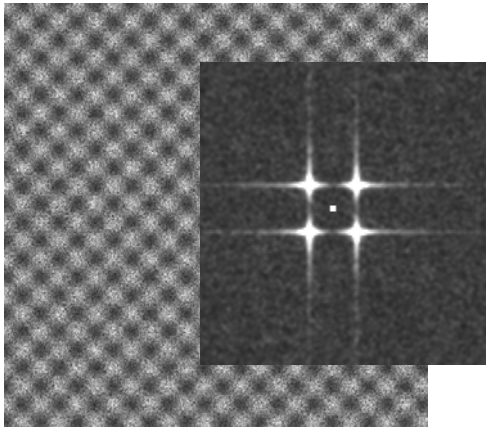


Image by Matt Chisholm,  
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Andy Lupini  
Aberration correction by Pete Nellist et  
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Nellist et al., Science 305, 2004.

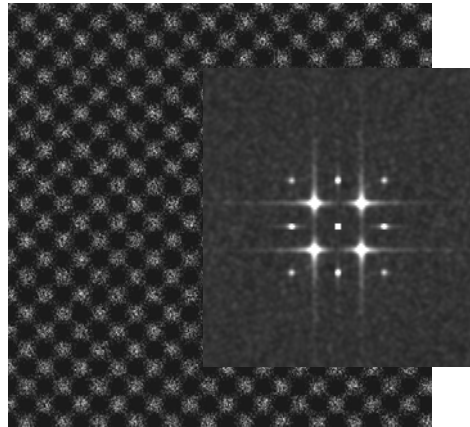
# Image artifacts

## *'clipping'*

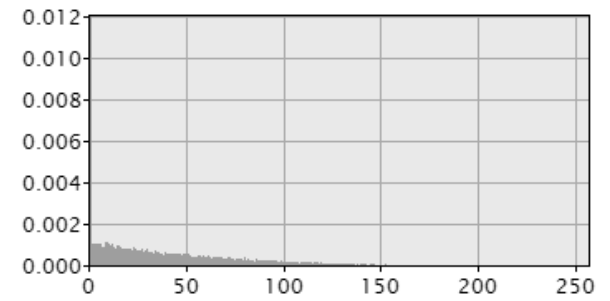
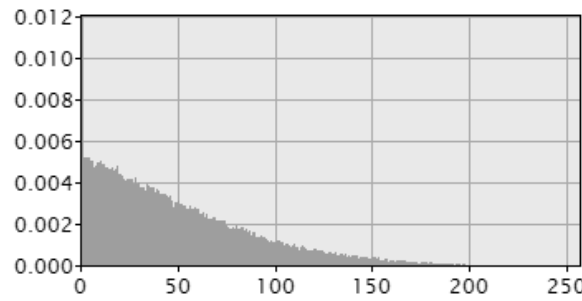
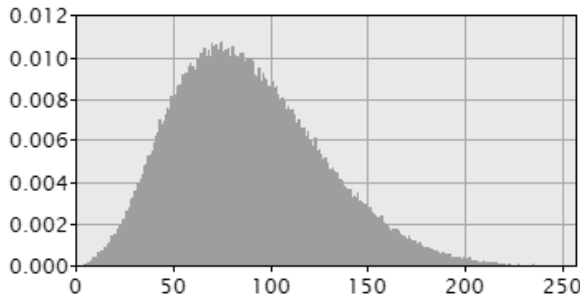
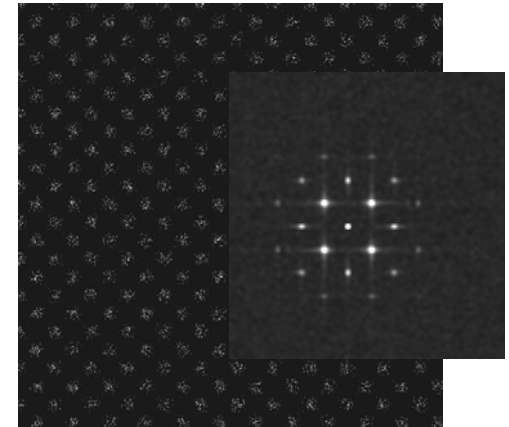
Unclipped



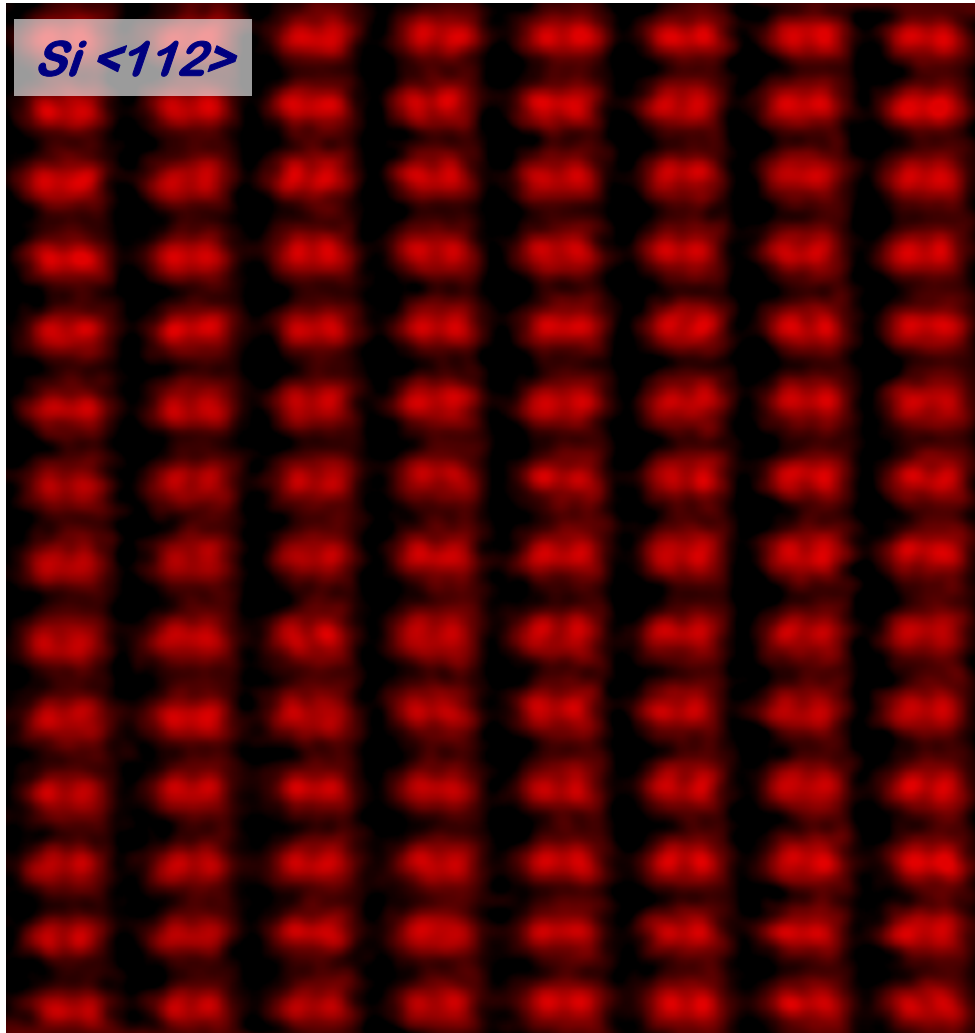
40% clipped



60% clipped



# Effect of $C_s$ correction



Direct resolution at 0.78 Å

Information transfer to 0.607 Å

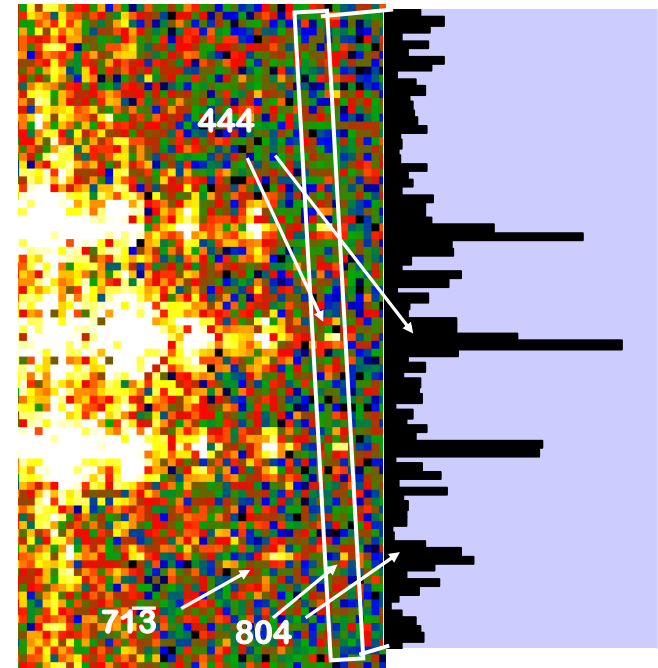


Image by Matt Chisholm,  
Processing by Albina Borisevich and  
Andy Lupini  
Aberration correction by Pete Nellist et  
al., Nion Co.  
Nellist et al., Science 305, 2004.